# Introduction to <u>Basics</u> of Quantum Mechanics for **Solar Engineering Physics** and **PV Power Plants**.

A set of solved example problems. From chapters 2, 3, 4, and 5 of the book titled 'Quantum Mechanics A Textbook for Undergraduates'. Mahesh C. Jain. PHI. 2007.

Entered by: Karl S. Bogha. Year: 2019. ME Studies, University of Canterbury. New Zealand. Any errors and omissions apologies in advance. Original resources were copyrighted material. This file is available for distribution. Quantum Mechanics A Textbook for Undergraduates. Mahesh C. Jain. PHI. Department of Physics, Hindu College University of Delhi. To Support Studies In: Modern Physics by S.N. Ghosal.

Purpose: <u>Quantum Mechanics for 'Power Plant Engineering' Studies.</u> Exercise by: K S Bogha.

| Constants:<br>h:= $6.63 \cdot 10^{-34}$ | J/s   | h is the 'quantum of energy'. |
|---|-------|-------------------------------|
| c≔3•10 <sup>8</sup>                     | ms^-1 |                               |
| $eV := 1.6 \cdot 10^{-19}$              | J     |                               |
| e≔1.6•10 <sup>-19</sup>                 | J     |                               |
|   |       |                               |

## Problem 2.1

Show that Planck's law reduces to Wein's law in the short wavelength limit and Rayleigh-Jean's law in the long wavelength limit.

## Solution:

See Chapter 3 Planck's Theory in Physics textbook by SN Ghoshal. Page 43. Both short and long wavelength solved. <u>OR other Modern Physics textbooks.</u>

#### Problem 2.2

Find the number of photons emitted per second by a 40W source of monochrome light of wavelength 6000 A (Angstrom - unit of length for measuring wavelength)

## Solution:

| nhv = E      | n is number of photons emitted per second  |
|--------------|--|
| n = E/(hv)   |  |
| n = E(lamda) | '(hc)                                      |
|              | $= 1/t_{1} \qquad (1 - t_{2} - 1 - t_{2})$ |

| V = | = c/ | (lam | da), | SO | 1/v | = ( | lamo | a/c |
|-----|------|------|------|----|-----|-----|------|-----|
|     |      |      |      |    |     |     |      |     |

| $\lambda := 6000 \cdot 10^{-10}$ | Angstrom |
|----------------------------------|----------|
| E≔40 Watts                       | 5        |
| $h = 6.63 \cdot 10^{-34}$        |          |
| $c := 3 \cdot 10^8$              |          |
|                                  |          |

$$n \coloneqq \frac{(E \cdot \lambda)}{(h \cdot c)} = 1.207 \cdot 10^{20}$$

 $n = 1.207 \cdot 10^{20}$  Number of photons per second. Ans.

|  | oemission occur if a photon of energy 3.8eV is incident on the surface?   |
|--|---|
| n yes, m   | nd in joules the maximum kinetic energy of the photoelectron.   |
| Solution:  |   |
|  | y of the photon is more than the work function of the surface, nergy), photoemission will occur:  |
| Work   | function of a metal (or material) is the minimum amount of energy required nove an electron from its surface.   |
| Work   | FunctionSurface = 3.2 eV  |
| Photo  | nEnergy:=3.8 eV   |
| KE_of  | photoelectron = PhotonEnergy – WorkFunctionSurface  |
| KE_of  | $f_{photoelectron} = 9.6 \cdot 10^{-20} \qquad \text{Ans.}$   |
| Problem  | 2.4   |
|  |   |
| I he work  | c function of a metal is 3.45 eV.   |
|  | c function of a metal is 3.45 eV.<br>he maximum wavelength of a photon that can eject an electron from the  |
| What is t  |   |
| What is t<br>metal.<br>Solution:<br>v_o is the   | he maximum wavelength of a photon that can eject an electron from the   |
| What is t<br>metal.<br>Solution:<br>v_o is the<br>h v_o = V  | he maximum wavelength of a photon that can eject an electron from the<br>e tresh hold frequency<br>WF (work function)   |
| What is t<br>metal.<br>Solution:<br>$v_o$ is the<br>$h v_o = v'$<br>$v_o = c/l$  | he maximum wavelength of a photon that can eject an electron from the<br>e tresh hold frequency<br>WF (work function)   |
| What is t<br>metal.<br>Solution:<br>$v_o$ is the<br>$h v_o = v$<br>$v_o = c/l$<br>now hc/la  | he maximum wavelength of a photon that can eject an electron from the<br>e tresh hold frequency<br>WF (work function)<br>amda_o   |
| What is t<br>metal.<br>Solution:<br>$v_o$ is the<br>$h v_o = V$<br>$v_o = c/l$<br>now hc/la<br>so lamda  | he maximum wavelength of a photon that can eject an electron from the<br>e tresh hold frequency<br>WF (work function)<br>amda_o<br>amda_o = WF  |
| What is t<br>metal.<br>Solution:<br>$v_o$ is the<br>$h v_o = V$<br>$v_o = c/I$<br>now hc/Ia<br>so lamda<br>WF := 3.4   | he maximum wavelength of a photon that can eject an electron from the<br>e tresh hold frequency<br>WF (work function)<br>amda_o<br>amda_o = WF<br>_o = hc/WF  |
| What is t<br>metal.<br>Solution:<br>$v_o is theh v_o = V$<br>$v_o = c/I$<br>now hc/Ia<br>so lamda<br>WF := 3.4<br>$\lambda \_:= \frac{(h)}{M}$                   | he maximum wavelength of a photon that can eject an electron from the<br>e tresh hold frequency<br>WF (work function)<br>amda_o<br>amda_o = WF<br>_o = hc/WF  |
| What is t<br>metal.<br>Solution:<br>$v_0$ is the<br>$h v_0 = V$<br>$v_0 = c/l$<br>now hc/la<br>so lamda<br>WF := 3.4<br>$\lambda _ := \frac{(h)}{M}$<br>Angstron | he maximum wavelength of a photon that can eject an electron from the<br>e tresh hold frequency<br>WF (work function)<br>amda_o<br>amda_o = WF<br>_o = hc/WF<br>$45 \cdot (1.6 \cdot 10^{-19})$<br>$\frac{\cdot c)}{/F} = 3.603 \cdot 10^{-7}$ meter  |
| What is t<br>metal.<br>Solution:<br>$v_0$ is the<br>$h v_0 = V$<br>$v_0 = c/l$<br>now hc/la<br>so lamda<br>WF := 3.4<br>$\lambda _ := \frac{(h)}{M}$<br>Angstror | he maximum wavelength of a photon that can eject an electron from the<br>te tresh hold frequency<br>WF (work function)<br>amda_o<br>amda_o = WF<br>_o = hc/WF<br>H5 $\cdot$ (1.6 $\cdot$ 10 <sup>-19</sup> )<br>$\frac{\cdot C}{/F} = 3.603 \cdot 10^{-7}$ meter<br>m_unit := 1 $\cdot$ 10 <sup>-10</sup> measure of Angstrom |

| A meta                     | l of wor             | c functio          | า 3.06         | eV is ill      | uminat   | ed bv  | liaht o | faw   | avelen | ath 30 | )00 Ar | nastorr | n. |
|----------------------------|----------------------|--------------------|----------------|----------------|----------|--------|---------|-------|--------|--------|--------|---------|----|
| Calcula                    |                      |                    |                |                |          | J      | 5       |       |        | J      |        | 3       |    |
| (a). tre                   | sh hold              | frequenc           | У              |                |          |        |         |       |        |        |        |         |    |
|                            |                      | energy o           |                | toeletre       | ons      |        |         |       |        |        |        |         |    |
| (c). sto                   | pping p              | otential           |                |                |          |        |         |       |        |        |        |         |    |
| <b>.</b>                   |                      |                    |                |                |          |        |         |       |        |        |        |         |    |
| Solutio                    | า:                   |                    |                |                |          |        |         |       |        |        |        |         |    |
| a).                        |                      |                    |                |                |          |        |         |       |        |        |        |         |    |
|                            |                      | hold fre           |                | су             |          |        |         |       |        |        |        |         |    |
|                            | -                    | ork funct          | ion)           |                |          |        |         |       |        |        |        |         |    |
| V_0 = \                    | NF/h                 |                    |                |                |          |        |         |       |        |        |        |         |    |
| \//E                       | $0 \sim 1$           | eV = 1             | 6 10           | <sup>−19</sup> |          |        |         |       |        |        |        |         |    |
| VVF := 3                   | .0 ev                | ev = 1             | .0•10          | ,              |          |        |         |       |        |        |        |         |    |
|                            | W/F                  |                    | 14             |                |          |        |         |       |        |        |        |         |    |
| V_0:                       | = <u> </u>           | 7.24•10            | ) ' ¯          | Hz. Ar         | IS.      |        |         |       |        |        |        |         |    |
|                            |                      |                    |                |                |          |        |         |       |        |        |        |         |    |
| b).                        |                      |                    |                |                |          |        |         |       |        |        |        |         |    |
| v = c/la                   | amda                 |                    |                |                |          |        |         |       |        |        |        |         |    |
| c:=0                       | clea                 | ar                 |                |                |          |        |         |       |        |        |        |         |    |
| c≔3•1                      | 0 <sup>8</sup> m/s   | 5                  |                |                |          |        |         |       |        |        |        |         |    |
| λ≔300                      | 0•10 <sup>-10</sup>  |                    |                |                |          |        |         |       |        |        |        |         |    |
| c_2 1                      | n <sup>8</sup>       | m/s spee           | d of I         | iaht           |          |        |         |       |        |        |        |         |    |
|                            |                      |                    |                |                |          |        |         |       |        |        |        |         |    |
| $V := \frac{C}{\lambda} =$ | 1 • 10 <sup>15</sup> | Hertz              | (1/s)          | freque         | ncy of   | incide | nt radi | ation | Ans.   |        |        |         |    |
| c).                        |                      |                    |                |                |          |        |         |       |        |        |        |         |    |
| Maximu                     | ım ener              | gy E = h           | (v - v <u></u> | _o)            |          |        |         |       |        |        |        |         |    |
| F                          | =h•(v-               | -v o)              |                |                |          |        |         |       |        |        |        |         |    |
|                            |                      |                    |                |                |          |        |         |       |        |        |        |         |    |
| E <sub>photon</sub>        | = 1.83•              | 10 <sup>-19</sup>  | J.             |                |          |        |         |       |        |        |        |         |    |
| Energy                     | in elect             | ron volt e         | eV:            |                |          |        |         |       |        |        |        |         |    |
| Enhator                    | $E_{pl}$             | $\frac{1}{V} = 1.$ | 144            | eV An          | S.       |        |         |       |        |        |        |         |    |
| -pnoton_                   | ev ·                 |                    |                | J. / 111       | <u> </u> |        |         |       |        |        |        |         |    |

d).  
Stopping potential Vo = Emax/e = (h/e)(v · v\_o)  
V<sub>0</sub> := 
$$\left(\frac{h}{e}\right) \cdot (v - v_o)$$
 V<sub>0</sub> = 1.144 Volt Ans.  
OR  
V<sub>0</sub> :=  $\frac{E_{photon}}{eV}$  = 1.144 Volt Ans.  
Problem 2.6  
Find the frequency of the light which ejects from a metal surface electrons, fully  
stopped by a retarding potential of 3V.  
The photoelectric effect begins in this metal at a frequency of 6x10^14 s^-1.  
Find the work function for this metal.  
Solution:  
Threshold frequency v\_0:=6.10<sup>14</sup>  
WF = h v\_0  
WF := h v\_0 = 3.978 \cdot 10<sup>-19</sup>  
WF  $= h v_0 = 3.978 \cdot 10^{-19}$   
WF  $= h v_0 = 3.978 \cdot 10^{-19}$   
WF  $= (eV_0 + h_v v_0)/h$   
retarding potential =  $3eV = eV = 1.6 \cdot 10^{-19}$   
 $eV_0 = 3 \cdot eV = 4.8 \cdot 10^{-19}$   
 $v_{=} \frac{(eV_0 + h_v v_0)}{(h)}$   
 $v = 1.324 \cdot 10^{15}$  Hz Ans.

|   | Sodium) is 2.3eV.<br><u>notoelectric emission</u> for light of wavelength 6800 Armstrong?                         |
|---|---|
| Solution:   |   |
| $\lambda \coloneqq 6800 \cdot Angstrom$                             | $\lambda = (6.8 \cdot 10^{-7}) \ m$ c := 3 \cdot 10 <sup>8</sup> $\frac{m}{s}$ h := 6.63 \cdot 10^{-34} J \cdot s |
| $eV = 1.6 \cdot 10^{-19}$ e   | $V := 1.6 \cdot 10^{-19} J$   |
| Energy of incident pho  | oton = hc/(lamda)   |
| Energy of incident pho  | oton = hc/lamda (ie at this wavelength)   |
| $E_{photon} := \frac{(h \cdot c)}{\lambda} = (2.9)$                 | $(25 \cdot 10^{-19}) J$   |
| $E_{photon_eV} := \frac{E_{photon}}{eV} =$                          | 1.828 eV Ans.   |
|   | oton 1.83 eV is less than 2.3 eV,<br>sion is possible with the given light (wavelength).                          |
| roblem 2.8  |   |
| 0   | 500 Angstrom is incident on two metals A and B.<br>Shotoelectrons if their work functions are 4.2 eV and 1.9 eV   |
| he work function of th  | ne metal is 0.1 eV.   |
| Solution:   |   |
| $\lambda := 3500 \cdot 10^{-10}$ Ang<br>h := 6.6 \cdot 10^{-34} J.s | strom<br>$c := 3 \cdot 10^8$ m/s $eV := 1.6 \cdot 10^{-19}$<br>$D^{-19}$ $E_eV := \frac{E}{eV} = 3.536$ $eV$      |
| $F = (h \cdot c) = 5.657 \cdot 10^{-5}$                             | $D^{-19}$ E_ev := <u>E</u> = 3.536 eV   |

Problem 2.9

Calculate the maximum kinetic energy of a photoelectron (in eV) emitted on shining light of wavelength 6.2 x 10<sup>-6</sup>m on a metal surface.

The work function of the metal surface is 0.1 eV.

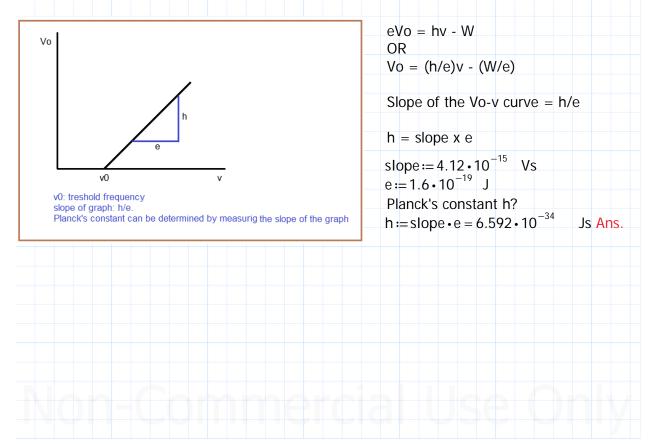
| Solution:                               |                                       |        |                              |                         |
|---|---------------------------------------|--------|------------------------------|-------------------------|
| $\lambda \coloneqq 6.2 \cdot 10^{-6}$ r | n W≔0.1                               | eV     | $h = 6.6 \cdot 10^{-34}$ J.s | c≔3•10 <sup>8</sup> m/s |
|   | c energy of a pho<br>= (hc/lamda) - W |        | is given by:                 |                         |
| Emov. (h•c)                             | $\frac{0}{V}$ - (W) = 0.1             | -V Ans |                              |                         |

Problem 2.10

In an experiment on photoelecric effect, the slope of the cut-off voltage versus frequency of incident light graph is found to be  $4.12 \times 10^{-15}$  Vs. Given  $e = 1.6 \times 10^{-19}$  C.

Estimate the value of Planck's constant.

Solution:



| Soluti | OD'  |
|--------|--|
|        |  |
|        | W_potassium := $2.26 \cdot eV = 3.616 \cdot 10^{-19}$ J (convert eV to Joules).  |
|        | Maximum kinetic energy (1/2) m v^2= hv - W<br>$hv = (1/2) m v^2 + W$<br>$v = ((1/2) m v^2 + W) / h$  |
|        | m_potassium := $(9 \cdot 10^{-31})$ mass of potassium electron kg  |
|        | max_speed_potassium_electron := (10 <sup>6</sup> ) m/s   |
|        | e_max_speed := $\left(\frac{1}{2}\right) \cdot (m_{\text{potassium}}) \cdot (max_{\text{speed}_{\text{potassium}_{\text{electron}}})^2$                            |
|        | $e_{max_speed} = 4.5 \cdot 10^{-19}$   |
|        | hv_potassium := e_max_speed + W_potassium = 8.116 • 10 <sup>-19</sup>  |
|        | v_frequency  |
|        | frequency_v:= $\frac{(hv_potassium)}{(h)} = 1.231 \cdot 10^{15}$ Hz Ans.   |
| Comi   | ment: An incandescent lamp frequency is 50 or 60 Hz, (just under 10^2 ie 100 Hz) his potassium electron's frequency is at 10^15 Hz. This is several millions times |

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| Problem 2.12 |  |  |  |
|--------------|--|--|--|
|              |  |  |  |
| (a).         |  |  |  |

A stopping potential <u>of 0.82 V is required</u> to stop the emission of photoelectrons from the surface of a metal by light of wavelength <u>4000 Angstrom</u>. For light of wavelength <u>3000</u> <u>Angstrom</u>, the stopping potential is <u>1.85 V</u>. Find the value of Planck's constant.

(b).

At stopping potential, if the wavelength of the incident light is kept fixed at 4000 Angstorm but the intensity of light is increased two times, will photoelectric current be obtained? Give reasons for your answer.

Solution:

a).

 $(hc/lamda-1) = eV1 + W \dots equation 1.$ 

 $(hc/lamda-2) = eV2 + W \dots equation 2.$ 

W is the same for equation 1 and 2 because its the same metal surface.

Subtracting 1 from 2

hc ( (1/lamda-2) - (1/lamda-1) ) = e(V2 - V1)

Rearrainging for h:

h = (e (V2 - V1)) / c ((1/lamda-1) - (1/lamda-2))

 $eV = 1.6 \cdot 10^{-19}$  V1 := 0.82 V2 := 1.85  $\lambda = 4000 \cdot 10^{-10}$   $\lambda = 3000 \cdot 10^{-10}$ 

c:=3.10<sup>8</sup>

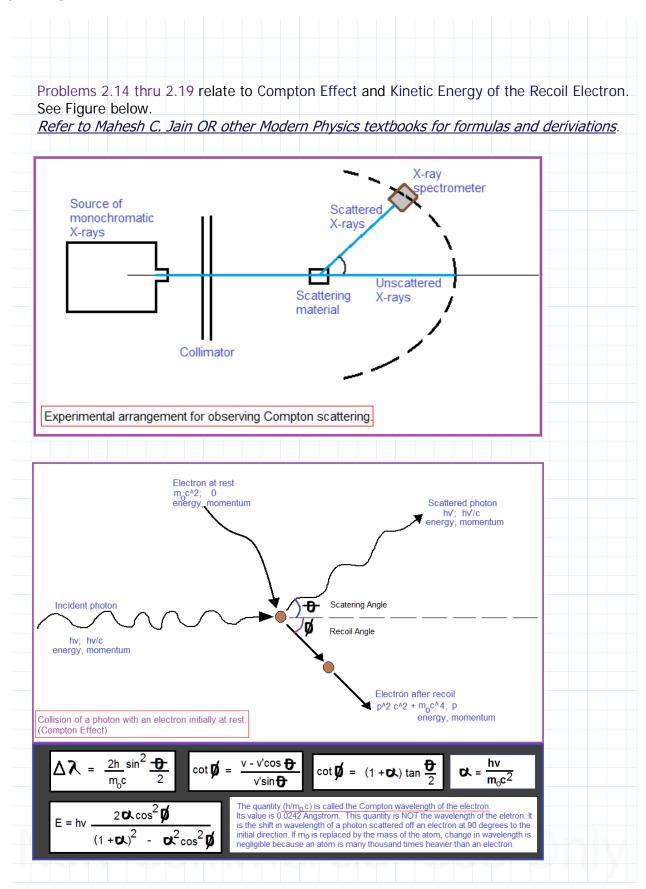
 $A := eV \cdot (V2 - V1) = 1.648 \cdot 10^{-19}$ 

 $\mathsf{B} := \mathsf{C} \cdot \left( \left( \frac{1}{\lambda 2} \right) - \left( \frac{1}{\lambda 1} \right) \right) = 2.5 \cdot 10^{14}$ 

 $h := \frac{A}{B} = 6.592 \cdot 10^{-34}$  Js. Ans.

b). No. When the intensity is increased the stopping potential does not increase because the stopping potential depends only on the wavelength of light NOT its intensity. Ans.

| Calcu<br>Work<br>Solu<br>W_( | late the       | elength<br>e photoel<br>n of Ces | lectric   | currer  | nt, assu            |         |                   |       |         |          |         |        | face. |
|------------------------------|----------------|----------------------------------|-----------|---------|---------------------|---------|-------------------|-------|---------|----------|---------|--------|-------|
| Vork<br>Solu<br>W_(          | functio        |                                  |           |         |                     | iming   |                   |       | officio | may a    |         |        |       |
| Solu<br>W_0                  |                |                                  | ium –     | 1.75    |                     | - 6 62  |                   |       |         | ncy of   | 0.5%    | •      |       |
| W_(                          | ition:         |                                  |           |         | CV, 11 -            | - 0.02  | X 10              | 54.5  | ,,,,    |          |         |        |       |
| W_(                          | Ition:         |                                  |           |         |                     |         |                   |       |         |          |         |        |       |
| Pov                          |                |                                  |           |         |                     |         |                   |       |         |          |         |        |       |
| Pow                          | cesium         | = 1.93                           | eV        |         |                     |         |                   |       |         |          |         |        |       |
| e := '                       | ior cos        | ium - 1.                         | $10^{-3}$ | mW      |                     |         |                   |       |         |          |         |        |       |
|                              | 1.6•10         | $0^{-19}$                        | J         |         |                     |         |                   |       |         |          |         |        |       |
| h:=                          | 6.62 • 10      | $0^{-34}$                        | J         |         |                     |         |                   |       |         |          |         |        |       |
| с := 1                       | $3 \cdot 10^8$ | -                                | m/s       |         |                     |         |                   |       |         |          |         |        |       |
| λ:=                          | 4560.1         | $0^{-10}$                        | 111/ 0    |         |                     |         |                   |       |         |          |         |        |       |
|                              |                | one phot                         |           |         | nc)/(lar            | nda)    |                   |       |         |          |         |        |       |
|                              | 0,5            |                                  |           |         |                     |         |                   |       |         |          |         |        |       |
| Ene                          | rav 1p         | hoton ≔                          | (h•c)     | = 4.3   | 55•10               | –19     | J                 |       |         |          |         |        |       |
| <b>_</b> 110                 | · 97_·P        | notoni                           | (λ)       |         | 00 10               |         | 0                 |       |         |          |         |        |       |
| Nur                          | nber of        | photons                          | incide    | ent on  | the su              | rface o | of ces            | ium p | er seco | ond      |         |        |       |
|                              |                |                                  |           |         |                     |         |                   |       |         |          |         |        |       |
| No                           | photon         | is≔ <u>Po</u><br>Ene             | wer_c     | esium   | _= 2.2              | 96•10   | ) <sup>15</sup>   |       |         |          |         |        |       |
|                              |                | Ene                              | rgy_1     | photo   | n                   |         |                   |       |         |          |         |        |       |
|                              |                |                                  |           |         |                     |         |                   |       |         |          |         |        |       |
| Sinc                         | e only         | 0.5% of                          | the pl    | notons  | incide              | nt on   | the ce            | esium | surfac  | e rele   | ase ele | ctrons | , the |
|                              | -              | electron                         |           |         |                     |         |                   |       |         |          |         |        |       |
|                              |                |                                  |           |         |                     |         |                   |       |         |          |         |        |       |
| No_                          | electro        | ns≔No                            | _phot     | ons•{-  | $\frac{0.5}{100} =$ | 1.148   | •10 <sup>13</sup> | ;     |         |          |         |        |       |
|                              |                |                                  |           | (       | 100/                |         |                   |       |         |          |         |        |       |
| Fina                         | illy we r      | reach to                         | where     | e we ca | an now              | calcu   | late t            | he ph | otoeleo | ctric cu | urrent, | ne:    |       |
|                              |                |                                  |           |         |                     |         |                   | 6     |         |          |         |        |       |
| Pho                          | toelect        | ric_curr                         | ent≔      | No_el   | ectron              | s•e='   | 1.837             | •10 ° | Amps    | . Ans.   |         |        |       |
|                              |                |                                  |           |         |                     |         |                   |       |         |          |         |        |       |
|                              |                |                                  |           |         |                     |         |                   |       |         |          |         |        |       |
|                              |                |                                  |           |         |                     |         |                   |       |         |          |         |        |       |
|                              |                |                                  |           |         |                     |         |                   |       |         |          |         |        |       |
|                              |                |                                  |           |         |                     |         |                   |       |         |          |         |        |       |
|                              |                |                                  |           |         |                     |         |                   |       |         |          |         |        |       |
|                              |                |                                  |           |         |                     |         |                   |       |         |          |         |        |       |
|                              |                |                                  |           |         |                     |         |                   |       |         |          |         |        |       |
|                              |                |                                  |           |         |                     |         |                   |       |         |          |         |        |       |
|                              |                |                                  |           |         |                     |         |                   |       |         |          |         |        |       |
|                              |                |                                  |           |         |                     |         |                   |       |         |          |         |        |       |



|   | wavelength 2.0 Angstrom are scattered from a carbon block.<br>ered photons are observed at <u>right angles</u> to the direction of the incident bean |
|---|--|
| Calculate:<br>a). The w                               | avelength of the scattered photon.   |
| b). The ei  | nergy of the recoil electron.  |
| c). The ar  | ngle at which the recoil electron appears.   |
|   | st mass of an electron m_o = 9.1 x 10 <sup>-31</sup> kg, c = 3 x 10 <sup>8</sup> m/s, and = 6.6 x 10 <sup>-34</sup> Js.                              |
| Solution:   |  |
| a).   |  |
|   | a 'are the wavelength of theincident and scattered photon, and the scattering angle, then  |
| delta lamo<br>λ≔2.0•1                                 | da = lamda' - lamda =(h/m_o c)(1 - cos (theta)).<br>0 <sup>-10</sup> Its already in Angstrom; ie 10^-10.   |
| $h = 6.6 \cdot 1$                                     | 0 <sup>-34</sup> J   |
| C:= 3 • 10  | m/s  |
| $m_0 := 9.1 \cdot \theta := 90 \ de$                  |  |
| $\Delta \lambda \coloneqq \left(\frac{h}{m_0}\right)$ | $\frac{1}{1 - \cos(\theta)} = 2.418 \cdot 10^{-12}$  |
| $\frac{\Delta\lambda}{10^{-10}} = 0.$                 | .024 Angstrom.   |
| lamda' =  | lamda + delta lamda = (h/m_o c)(1 - cos (theta)).  |
| $\lambda \coloneqq \lambda + \Delta$                  | $\lambda = 2.024 \cdot 10^{-10}$ Ans.  |
| b). Negleo  | cting the binding energy of the electron, its recoil energy is given by  |
|   | v - v') where v is the frequency<br>: ( (1/lamda) - (1/lamda') )   |
|   | •c) • $\left(\left(\frac{1}{\lambda}\right) - \left(\frac{1}{\lambda}\right)\right) = 1.182 \cdot 10^{-17}$ J Ans.                                   |

Exercise by: K S Bogha.

c). The angle theta at which the recoil electron appears is given by  

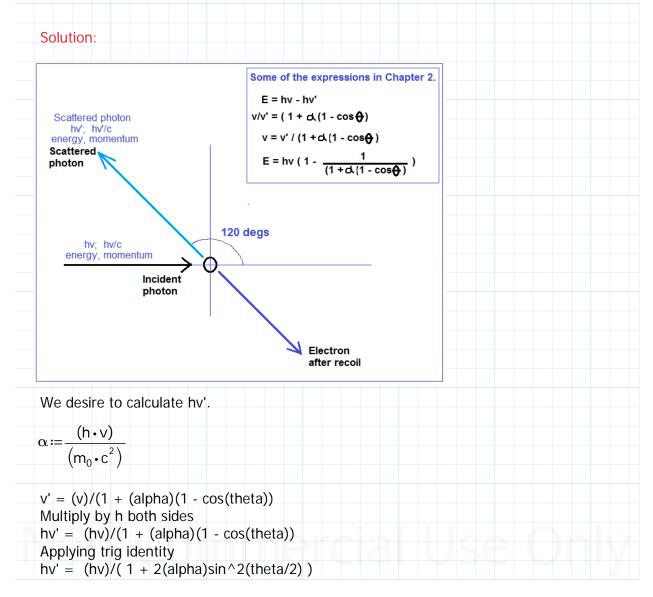
$$\cot (\text{phi}) = (1/\sin(\text{theta}) ((v/v') - \cos(\text{theta})))$$

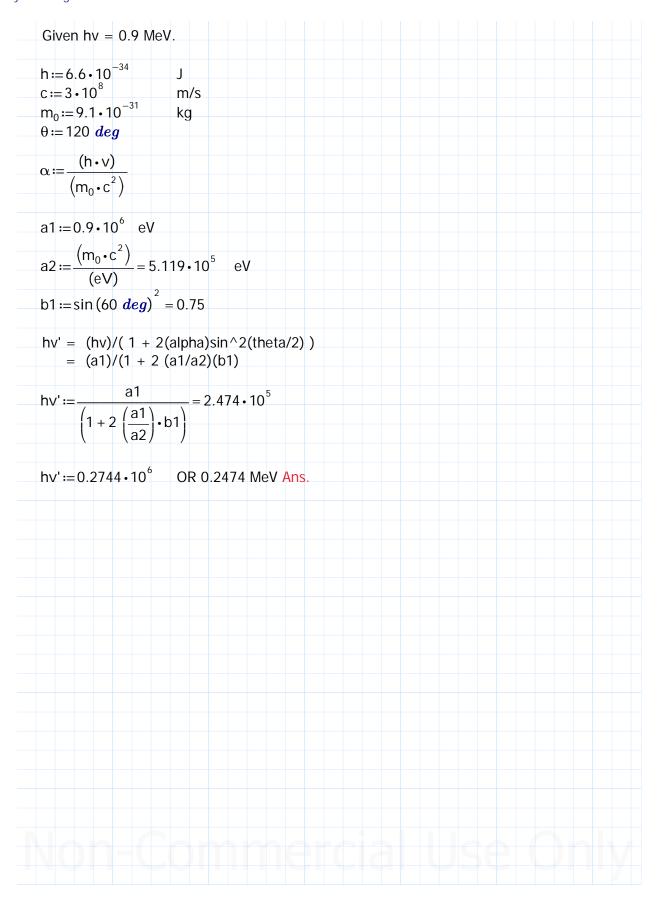
$$= (1/\sin(\text{theta}) ((\text{lamda'/lamda}) - \cos(\text{theta}))$$

$$\cot (\phi) := \left(\frac{1}{\sin(\theta)}\right) \cdot \left(\left(\frac{\lambda}{\lambda}\right) - \cos(\theta)\right) = 1.012$$

$$\phi := \arctan(1.012) = 44.658 \text{ deg} \text{ Ans.}$$
Problem 2.15

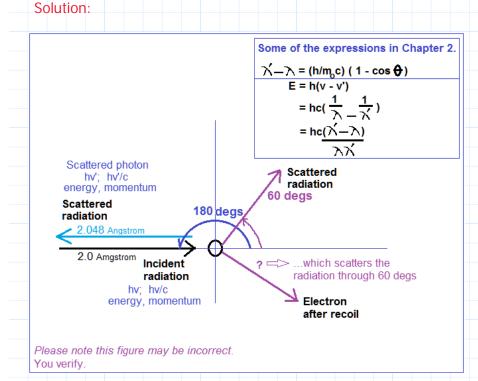
A photon of energy 0.9 MeV is scattered through 120 degs by a free electron. Calculate the energy of the scattered photon.





| In a Compton scattering experiment, the incident radiation has wavelength 2.000 Angstrom |
|--|
| while the wavelength of the radiation scattered through 180 degs is 2.048 Angstrom.      |
| Calculate:   |

- a). the wavelength of the radiation scattered at an angle of 60 degs to the direction of incidence
- b). energy of the recoil electron which scatters the radiation through 60 degs.



 $(lamda)' - (lamda) = (h/m_o c)(1 - cos(theta))$ 

We know lamda and theta, so we can calculate (h/m-o c), here theta is 180 degs that is the radiation at 180 degs. At 180 degs we have a specific radiation 2.048, and at 0 deg (incident) we have a specific radiation 2.048. Solving for (h/m\_o c) then we proceed to 0 degs incidence and 60 degs scattered radiation.

$$\lambda' = \text{mi} \ \mathbf{n} = \mathbf{u} (2 \cdot 0 \cdot 48 - 2.0) \cdot 10^{-10} = 4.8 \cdot 10^{-12}$$

 $a1 := (1 - \cos(180 \ deg)) = 2$ 

Problem 2.16

h\_div\_m\_c := 
$$\frac{\lambda'}{2}$$
 mi n u s.4 -10<sup>-12</sup> m

a). When theta = 60 degs to the incident radiation (lamda)' = (lamda) + (h/m\_o c)(1 - cos(theta)) by rearranging

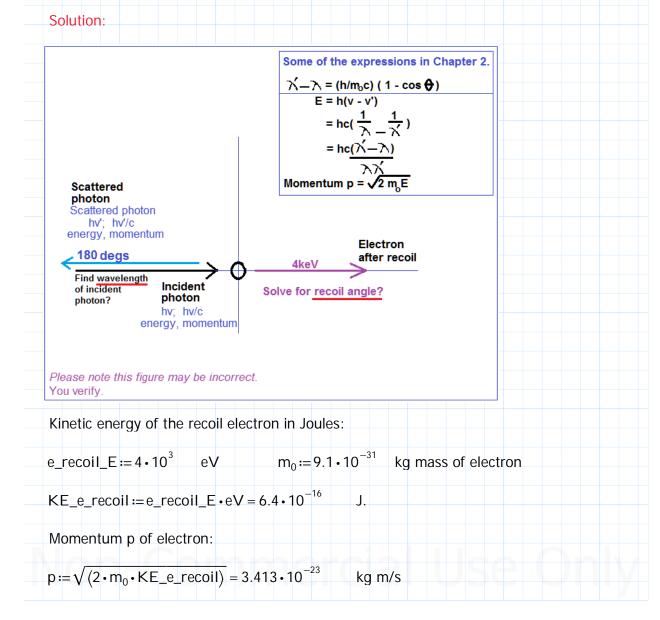
 $\lambda \coloneqq 2.000 \cdot 10^{-10} + (h_div_m_0c) \cdot (1 - \cos(60 \ deg)) = 2.012 \cdot 10^{-10}$  m Ans. OR Ans is 2.012 Angstrom. Quantum Mechanics A Textbook for Undergraduates. Mahesh C. Jain. PHI. Department of Physics, Hindu College University of Delhi. To Support Studies In: Modern Physics by S.N. Ghosal.

Purpose: <u>Quantum Mechanics for 'Power Plant Engineering' Studies.</u> Exercise by: K S Bogha.

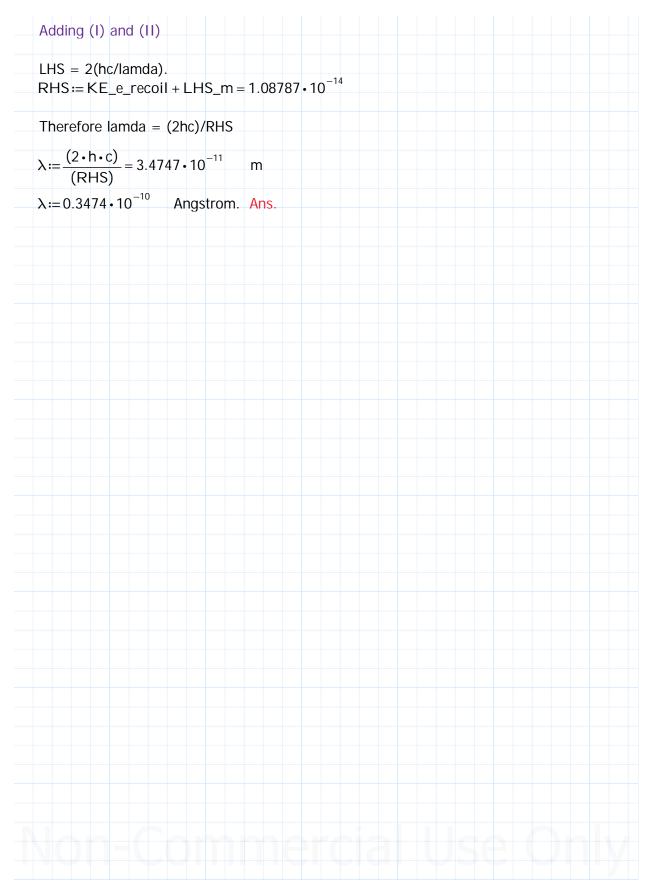
| b). We know both the wavelenghts; lamda' and lamda. we can plue these in to the hv - hv' expression where we substitute lamda f   |  |
|---|--|
| v. Where $v = 1/lamda$ .  |  |
| Therefore, hv - hv' = hc ((1/lamda') - (1/lamda))   |  |
| E_recoil_60degs := $(\mathbf{h} \cdot \mathbf{c}) \cdot \left(\left(\frac{1}{\lambda}\right) - \left(\frac{1}{\lambda}\right)\right) = 5.905 \cdot 10^{-18} \text{ J Ans.}$ |  |
| Problem 2.17  |  |

In a Compton scattering experiment, the X-ray photon is scattered at <u>an angle of 180 degs</u> and the electron recoils with an energy of 4keV.

Calculate the wavelength of the incident photon.



| Conservation of e<br>(hc/lamda) -  |   |  |
|--|---|--|
|  | (hc/lamda') =   | = KE of recoil electron = 6.4*10^-16 J(I)  |
| Exactly where is t   | the recoil elec   | ctron's direction?   |
| Use the cot(phi)   | = (1 + alpha)   | ) tan (theta/2) tp solve for angle psi.  |
| h≔6.3•10 <sup>-34</sup>  | J   |  |
| c ≔ 3 • 10 <sup>8</sup>  | m/s   |  |
| $m_0 := 9.1 \cdot 10^{-31}$  | kg  |  |
| $h := 6.3 \cdot 10^{-34}$<br>$c := 3 \cdot 10^{8}$<br>$m_{0} := 9.1 \cdot 10^{-31}$<br>$\theta := 180 \ deg$ | $\theta \mathbf{l} = \frac{\theta}{2} = 90$                       | deg  |
| $\alpha := \frac{(\mathbf{h} \cdot \mathbf{c})}{(\mathbf{m}_0 \cdot \mathbf{c}^2)} = 2.$                     | 308 • 10 <sup>-12</sup>   |  |
| a1 := $(1.0 + \alpha)$   | a1 = 1  | Approximately equal 1 the decimal value 2.418x10^-12is almost 0.   |
| $a2 := tan(\theta) = 1.$   | 633•10 <sup>16</sup>  |  |
| Therefore a1 x a2  | 2 = 1.0 appro   | oximately.   |
| conclude the ang<br>electron, the scat<br>recoil electron is   | e values of a<br>le phi is 0 deg<br>tered photon<br>in the same c | Not valid.<br>cot(1), a discussion or inspection may lead to<br>g. The angle phi is the angle of the recoil<br>is 180 deg ie the angle theta. So we say the<br>lirection as the incident photon - 0 degs.<br>egrees apart. See figure above. |
| Applying the <u>law</u><br>recoil + scatt  | of conservation   | on of momentum in the direction of the incident as positi<br>= incident:   |
| p cos(phi) + (h/la   | amda')cos(the   | eta) = h/(lamda)Eq 2.17-1  |
| b1 := cos (0 <i>deg</i> ) =<br>p c o : <b>s p</b> • b1 = 3.  |   | calculated p prior, therefore $p \times (cos(0) = p \times 1)$   |
| $b2 \coloneqq \cos(180 \ deg$  | g) = −1 as e  | expected -ve sign since scattered is in opposite direction ncident.  |
| Now after rearran<br>(h/lamda) + (h/la<br>Multiply both side<br>(hc/lamda) + (hc                             | amda') = 3.41<br>es by cthis                                      | 13*10^-23<br>now gives the conservation of momentum expression   |
| $LHS_m = p \cdot c = 1$  | .024 • 10 <sup>-14</sup>  |  |

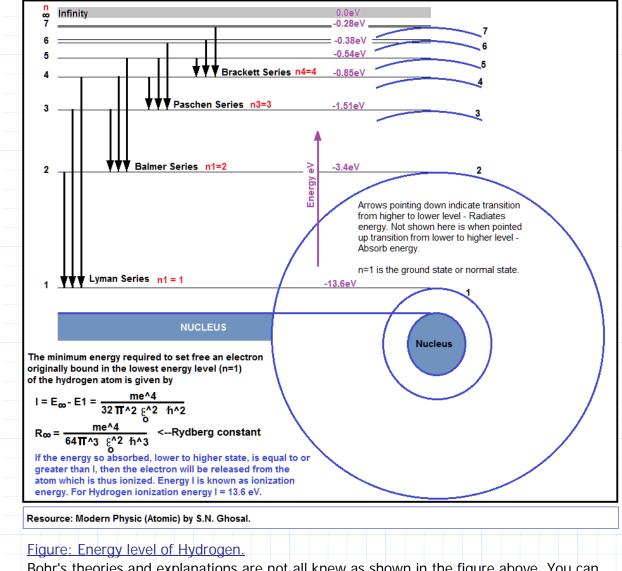


|    | . What is the maximum kinetic energy that can be imparted to a free electron by a photon of initial frequency v?  |
|----|---|
| b) | . Is it possible for the photon to transfer all its energy to the electron?   |
| So | plution:  |
|    | a). Kinetic energy of a recoil electron is  |
|    | E = (hv) <u>((alpha)(1 - cos(theta))</u><br>1 + (alpha)(1 - cos(theta))   |
|    | What is the impact of cos(theta)? If angle theta is 180 degs the cos(180 deg) = -1.<br>This value -1 increases the value of E it makes<br>E maximum.  |
|    | Emax when theta = 180 degs. Correct.  |
|    | Emax = (hv) <u>((alpha)(1 - (-1))</u><br>1+ (alpha)(1 - (-1))   |
|    | Emax = (hv) <u>((alpha)(2)</u><br>1+ (alpha)(2)   |
|    | $Emax = (hv)(\underline{2(alpha)}) Ans.$<br>1+ 2(alpha)   |
|    | b).   |
|    | For any real value alpha takes the RHS's left most term (2alpha/(1+2 alpha) is<br>less than 1. Which makes the 2nd term in the RHS term less than 1.0.<br>Hence, Emax < hv since hv is multiplied by the factor less than 1.0.                    |
|    | Emax < hv.<br>Since the maximum energy on the LHS is less than hv, the free electron cannot receiv<br>the maximum energy from the photon i.e. hv rather less than hv.   |
|    | So, a photon cannot transfer all its energy to a free electron. Thanks! Ans.  |
|    | Note: The Compton effect (delta)lamda = (h/m_o c) (1 - cos(theta)) here too we see when<br>angle theta is 180 degs, cos(180) = -1, then (1 - cos(180)) = 1 - (-1) = 2.<br>(delta) lamda = lamda' - lamda = 2(h/m_o c). The maximum Compton shift. |

|                                | ngle for symmetric scattering (i.e. theta = phi) at this energy.          |
|--------------------------------|---|
| b). What is th                 | e energy of the scattered photon for this case?                           |
| Solution:                      |   |
| a).                            | Scattered<br>photon   |
|                                | Condition:<br>1. Ø = <del>0</del>   |
| br. b                          |   |
| hv; hv<br>energy, m<br>Incid   | $rac{rac}{rac}$   |
| radia                          | tion y  |
|                                |   |
|                                | Electron<br>after recoil  |
| Relation betw                  | een theta and psi:  |
| cot (psi) = ( 1                | + alpha )( tan (theta/2) )  |
| Set angle phi                  | = angle theta.  |
| Rewriting in s                 | in and cos terms, and substitute angle pis for theta.                     |
| cos (theta)/sir                | n(theta) = (1 + alpha) (sin (theta/2)/cos(theta/2))                       |
| We want to so                  | olve for angle theta.   |
| The physicist<br>arranging to: | Mahesh C. Jain reduced the expression above using trig identities and re- |
| Sin(theta/2)                   | = 1 / (Sqrt(2(alpha + 2)you may verify this yourself                      |
|                                |   |

| alpha = hv/m_0 o  | c^2                                     |       |         |             |            |   |  |
|---|---|-------|---------|-------------|------------|---|--|
| Numerator = $1.02$<br>Denominator = h                     | 2 MeV<br>alf the numerator              | = 0.5 | 1 MeV   |             |            |   |  |
|   |   | 0.0   |         |             |            |   |  |
| $\alpha := \frac{1.02}{0.51} = 2$                         |   |       |         |             |            |   |  |
| - ' (+ + ')   | 1                                       |       |         |             |            |   |  |
| sin (theta_div_2)   | $:= \frac{1}{\sqrt{(2 \cdot (2 + 2))}}$ |       |         |             |            |   |  |
| sin (theta_div_2)   |   |       |         |             |            |   |  |
|   | √ <u>(8)</u>                            |       |         |             |            |   |  |
| θ≔2•(asin (0.35⁄  | 4)) = 41.464 <i>deg</i>                 | Ans.  | You may | verify this | s yourself | ; |  |
| b).   |   |       |         |             |            |   |  |
|   |   |       |         |             |            |   |  |
| Energy of the sca   | ittered photon:                         |       |         |             |            |   |  |
| hv' = hv / (1 + al  |   |       |         |             |            |   |  |
|   | ipna(1 - cos(theta)                     | ))    |         |             |            |   |  |
|   |   |       | W Ans   |             |            |   |  |
|   | $\frac{2}{\cos(\theta))} = 0.679$       |       | eV Ans. |             |            |   |  |
|   |   |       | eV Ans. |             |            |   |  |
| $hv' := \frac{(1.02)}{(1 + \alpha \cdot (1 - 1))}$        |   |       | eV Ans. |             |            |   |  |
| $hv' := \frac{(1.02)}{(1 + \alpha \cdot (1 - 1))}$        |   |       | eV Ans. |             |            |   |  |
| $hv' \coloneqq \frac{(1.02)}{(1 + \alpha \cdot (1 - 1))}$ |   |       | eV Ans. |             |            |   |  |
| $hv' := \frac{(1.02)}{(1 + \alpha \cdot (1 - 1))}$        |   |       | eV Ans. |             |            |   |  |
| $hv' \coloneqq \frac{(1.02)}{(1 + \alpha \cdot (1 - 1))}$ |   |       | eV Ans. |             |            |   |  |
| $hv' \coloneqq \frac{(1.02)}{(1 + \alpha \cdot (1 - 1))}$ |   |       | eV Ans. |             |            |   |  |
| $hv' := \frac{(1.02)}{(1 + \alpha \cdot (1 - 1))}$        |   |       | eV Ans. |             |            |   |  |
| $hv' := \frac{(1.02)}{(1 + \alpha \cdot (1 - 1))}$        | $\frac{2}{\cos(\theta))} = 0.679$       |       | eV Ans. |             |            |   |  |

> Bohr's model is essential for understanding the continuing chapters in modern/atomic physics. Loosely though there were later found flaws and replaced by quantum mechanic's explanation(s). It is essential for the student to understand Bohr's atomic model before progressing to more suitable theories. Bohr proposed 3 postulates which were in contradiction to laws of classical mechanics and electromagnetic theories. This subject matter may be found in any current UG modern physics textbook.Calculations are good, its just the assumption or background on postulates is off.



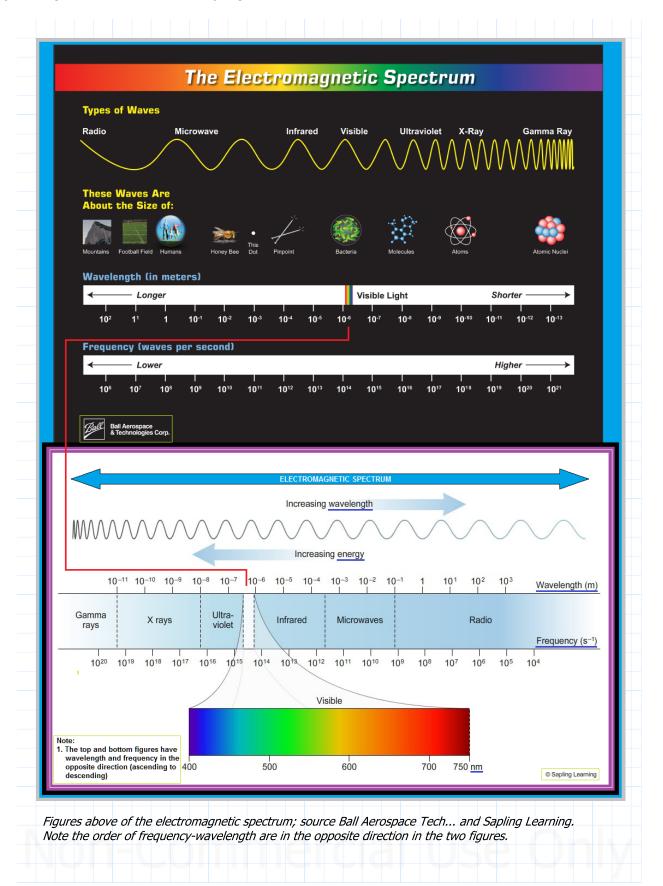
Bohr's theories and explanations are not all knew as shown in the figure above. You can related to them. You should have some understanding of the equations in the figure above, these would be found in most UG modern physics or quantum mechanics textbook.

The contents of the subject matter here leads to or generated the electromagnetic specturm. The spectrum we have seen in many ocassions in Physics, and Electrical Engineering courses. The parameters here are wavelength, and frequency. Part I we had Plank's contant h, wavelength (lamda), and frequency (v), eventually leading to the spectrum.

Quantum Mechanics A Textbook for Undergraduates. Mahesh C. Jain. PHI. Department of Physics, Hindu College University of Delhi. To Support Studies In: Modern Physics by S.N. Ghosal.

Purpose: Quantum Mechanics for 'Power Plant Engineering' Studies.

Exercise by: K S Bogha. Part II Structure of the Hydrogen Atom and The Bohr Model.



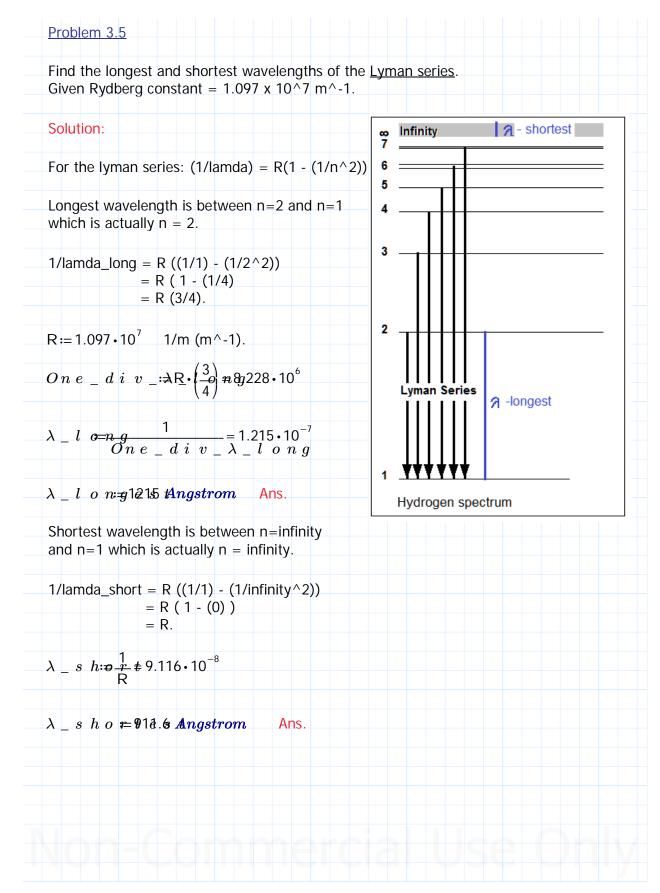
| $h = 6.63 \cdot 10^{-34}$   | Js   |  |
|---|--|--|
| $c := 3 \cdot 10^8$   | m/s  |  |
| $eV := 1.6 \cdot 10^{-19}$  |  |  |
| $ev := 1.6 \cdot 10$<br>$e := 1.6 \cdot 10^{-19}$   | J  |  |
| e≔ 1.6 • 10   | J  |  |
| Problem 3.1   |  |  |
|   | in exited hydrogen atom is -3.4eV.   |  |
| Calculate the ang   | ngular momentum of the electron according to Bohr theory.  |  |
| Solution:   |  |  |
| Brief notes:  |  |  |
| Bohr's radius of  | <sup>c</sup> Hydrogen a_0 = ( 4(pi)(epsilon_0)(h/2pi) )/ ( me^2) ).  |  |
| Energy in terms   | s of Bohr's radius a_0, En = -( (2pi/h) Z^2 ) / ( 2 m a_0 n^2 ).   |  |
|   | En = -13.6 Z^2 / n^2 eV  |  |
|   | E I I = -I J J U Z Z / I I Z E V   |  |
| for Hydrogen n  |  |  |
|   | = 1, and Z = 1, therefore En1_Hydrogen = -13.6 eV.<br>o other series relative to n and Z.  |  |
| similarly for the   | = 1, and Z = 1, therefore En1_Hydrogen = -13.6 eV.   |  |
| similarly for the<br>Angular moment   | = 1, and Z = 1, therefore En1_Hydrogen = -13.6 eV.<br>o other series relative to n and Z.<br>ntum = (nh/2 pi).   |  |
| similarly for the<br>Angular moment   | = 1, and Z = 1, therefore En1_Hydrogen = -13.6 eV.<br>o other series relative to n and Z.<br>htum = (nh/2 pi).<br>the hydrogen atom in the nth orbit : En = -13.6 /n^2 eV.   |  |
| similarly for the<br>Angular moment   | = 1, and Z = 1, therefore En1_Hydrogen = -13.6 eV.<br>o other series relative to n and Z.<br>ntum = (nh/2 pi).   |  |
| <i>similarly for the</i><br><i>Angular moment</i><br>Energy level of t  | = 1, and Z = 1, therefore En1_Hydrogen = -13.6 eV.<br>o other series relative to n and Z.<br>ntum = (nh/2 pi).<br>the hydrogen atom in the nth orbit : En = -13.6 /n^2 eV.<br>$n^2 = -13.6 \text{ eV}/(-3.4) \text{ eV}$   |  |
| <i>similarly for the</i><br><i>Angular moment</i><br>Energy level of t  | = 1, and Z = 1, therefore En1_Hydrogen = -13.6 eV.<br>o other series relative to n and Z.<br>ntum = (nh/2 pi).<br>the hydrogen atom in the nth orbit : En = -13.6 /n^2 eV.<br>$n^2 = -13.6 \text{ eV}/(-3.4) \text{ eV}$   |  |
| similarly for the<br>Angular moment<br>Energy level of t<br>n_squared := $\frac{-1}{-3}$                            | = 1, and Z = 1, therefore En1_Hydrogen = -13.6 eV.<br>o other series relative to n and Z.<br>ntum = (nh/2 pi).<br>the hydrogen atom in the nth orbit : En = -13.6 /n^2 eV.<br>$n^2 = -13.6 \text{ eV}/(-3.4) \text{ eV}$<br>13.6 = 4   |  |
| similarly for the<br>Angular moment<br>Energy level of t<br>n_squared := $\frac{-1}{-3}$                            | = 1, and Z = 1, therefore En1_Hydrogen = -13.6 eV.<br>o other series relative to n and Z.<br>ntum = (nh/2 pi).<br>the hydrogen atom in the nth orbit : En = -13.6 /n^2 eV.<br>$n^2 = -13.6 \text{ eV}/(-3.4) \text{ eV}$<br>13.6 = 4   |  |
| similarly for the<br>Angular moment<br>Energy level of t<br>n_squared := $\frac{-1}{-3}$<br>n := $\sqrt{n_squared}$ | = 1, and Z = 1, therefore En1_Hydrogen = -13.6 eV.<br>o other series relative to n and Z.<br>ntum = (nh/2 pi).<br>the hydrogen atom in the nth orbit : En = -13.6 /n^2 eV.<br>$n^2 = -13.6 \text{ eV}/(-3.4) \text{ eV}$<br>$\frac{13.6}{3.4} = 4$<br>$\frac{13.6}{3.4} = 2$ |  |
| similarly for the<br>Angular moment<br>Energy level of t<br>n_squared := $\frac{-1}{-3}$<br>n := $\sqrt{n_squared}$ | = 1, and Z = 1, therefore En1_Hydrogen = -13.6 eV.<br>o other series relative to n and Z.<br>ntum = (nh/2 pi).<br>the hydrogen atom in the nth orbit : En = -13.6 /n^2 eV.<br>$n^2 = -13.6 \text{ eV}/(-3.4) \text{ eV}$<br>13.6 = 4   |  |
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| similarly for the<br>Angular moment<br>Energy level of t<br>n_squared := $\frac{-1}{-3}$<br>n := $\sqrt{n_squared}$ | = 1, and Z = 1, therefore En1_Hydrogen = -13.6 eV.<br>o other series relative to n and Z.<br>ntum = (nh/2 pi).<br>the hydrogen atom in the nth orbit : En = -13.6 /n^2 eV.<br>$n^2 = -13.6 \text{ eV}/(-3.4) \text{ eV}$<br>$\frac{13.6}{3.4} = 4$<br>$\frac{13.6}{3.4} = 2$ |  |
| similarly for the<br>Angular moment<br>Energy level of t<br>n_squared := $\frac{-1}{-3}$<br>n := $\sqrt{n_squared}$ | = 1, and Z = 1, therefore En1_Hydrogen = -13.6 eV.<br>o other series relative to n and Z.<br>ntum = (nh/2 pi).<br>the hydrogen atom in the nth orbit : En = -13.6 /n^2 eV.<br>$n^2 = -13.6 \text{ eV}/(-3.4) \text{ eV}$<br>$\frac{13.6}{3.4} = 4$<br>$\frac{13.6}{3.4} = 2$ |  |
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| similarly for the<br>Angular moment<br>Energy level of t<br>n_squared := $\frac{-1}{-3}$<br>n := $\sqrt{n_squared}$ | = 1, and Z = 1, therefore En1_Hydrogen = -13.6 eV.<br>o other series relative to n and Z.<br>ntum = (nh/2 pi).<br>the hydrogen atom in the nth orbit : En = -13.6 /n^2 eV.<br>$n^2 = -13.6 \text{ eV}/(-3.4) \text{ eV}$<br>$\frac{13.6}{3.4} = 4$<br>$\frac{13.6}{3.4} = 2$ |  |
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| similarly for the<br>Angular moment<br>Energy level of t<br>n_squared := $\frac{-1}{-3}$<br>n := $\sqrt{n_squared}$ | = 1, and Z = 1, therefore En1_Hydrogen = -13.6 eV.<br>o other series relative to n and Z.<br>ntum = (nh/2 pi).<br>the hydrogen atom in the nth orbit : En = -13.6 /n^2 eV.<br>$n^2 = -13.6 \text{ eV}/(-3.4) \text{ eV}$<br>$\frac{13.6}{3.4} = 4$<br>$\frac{13.6}{3.4} = 2$ |  |

|                                    | energy of the photon emitted in the transition from $n=4$ to $n=2$ .   |
|------------------------------------|--|
| Solution:                          |  |
|                                    | on is falling from $n=4$ to $n=2$ , higher energy to lower energy, it is radiating/<br>energy. Initial state $ni = 4$ (higher energy), final state $nf = 2$ (lower energy).  |
| ni≔4 i                             | mitted = E_ground state x ( (1/ni-lower Energy^2) - (1/nf-higher Energy^2) )<br>nitial state - higher energy<br>final state - lower energy   |
| E <sub>1_H</sub> ≔—1               | 13.6 eV  |
| E <sub>4_2</sub> ≔E <sub>1_</sub>  | $_{-H} \cdot \left( \left( \frac{1}{ni^2} \right) - \left( \frac{1}{nf^2} \right) \right) = 2.55 \text{ eV } + \text{ve sign radiated or emitted energy Ans.}$   |
| The atom<br>The atom<br>state 2 to | rse the direction.<br>transitions from n = 2 to n =4.<br>will absorb energy because its travelling from lower energy<br>higher energy state 4.<br>reverse the order of the expression for ni and nf.   |
| E <sub>2_4</sub> :=E <sub>1_</sub> | $-H \cdot \left( \left( \frac{1}{nf^2} \right) - \left( \frac{1}{ni^2} \right) \right) = -2.55  \text{eV -ve sign energy absorbed Ans.}$   |
| Comment                            | <ul> <li>S: You verify on the sign convention.</li> <li>Its not exactly like Kirchchoff's Law in electric circuit's convention.</li> <li>Lets say if we set the convention, radiated outward is +ve sign,</li> <li>then absorbed would be -ve sign. Provided the ni and nf line up with the order of energy direction movement.</li> </ul> |
|                                    |  |
|                                    |  |

|  | a line of Balm<br>n = 2 (energy                     | er series is obta<br>v = -3.4eV).          | nined from the  | transition $n = 3$ | 3 (energy =      |
|--|---|--|---|--------------------|------------------|
| Calculate t  | ne wavelength                                       | n for this line.                           |   |                    |                  |
| Solution:  |   |  |   |                    |                  |
|  |   | ord 'line' used ir<br>m laboratory ex      |   | Its just like the  | e spectrum chart |
| General ex   |   |  |   |                    |                  |
| E2 > E1 er   | ergy level n=                                       | 2 is higher than                           | า 1   |                    |                  |
| hv = E2 - I  |   |  |   |                    |                  |
| frequency  | / = c/lamda   |  |   |                    |                  |
| (hc/lamda)   | = E2 - E1   |  |   |                    |                  |
| wavelength   | n lamda = hc /                                      | / (E2 - E1)                                |   |                    |                  |
| In our prot  | lem we go fro                                       | om E3 to E2 (hig                           | gher to lower -   | radiating energ    | av)              |
|  |   |  |   | 0 0                |                  |
|  |   |  |   |                    |                  |
| $E_3 := -1.5$<br>$E_2 := -3.4$                             | eV<br>eV  |  |   |                    |                  |
| E2≔-3.4  | eV  | ( 5 to 10 <sup>-7</sup>                    |   |                    |                  |
| E2≔-3.4  | eV  | $\frac{1}{\sqrt{1}} = 6.543 \cdot 10^{-7}$ | m.  |                    |                  |
| $E2 \coloneqq -3.4$ $\lambda \_ H_3 \coloneqq$             | eV<br>(h∙c)<br>(E3−E2)∙e\                           | $\frac{1}{\sqrt{2}} = 6.543 \cdot 10^{-7}$ | m.  |                    |                  |
| $E2 := -3.4$ $\lambda = H_{3} := 3$ $\lambda = H_{3} := 3$ | eV<br>(h∙c)<br>(E3−E2)∙e\<br>6543•10 <sup>-10</sup> |  | m.  |                    |                  |
| $E2 := -3.4$ $\lambda = H_{3} := 3$ $\lambda = H_{3} := 3$ | eV<br>(h∙c)<br>(E3−E2)∙e\                           |  | m.  |                    |                  |
| $E2 := -3.4$ $\lambda = H_{3} := 3$ $\lambda = H_{3} := 3$ | eV<br>(h∙c)<br>(E3−E2)∙e\<br>6543•10 <sup>-10</sup> |  | m.  |                    |                  |
| $E2 := -3.4$ $\lambda = H_{3} := 3$ $\lambda = H_{3} := 3$ | eV<br>(h∙c)<br>(E3−E2)∙e\<br>6543•10 <sup>-10</sup> |  | m.  |                    |                  |
| $E2 := -3.4$ $\lambda = H_{3} := 3$ $\lambda = H_{3} := 3$ | eV<br>(h∙c)<br>(E3−E2)∙e\<br>6543•10 <sup>-10</sup> |  | m.  |                    |                  |
| $E2 := -3.4$ $\lambda = H_{3} := 3$ $\lambda = H_{3} := 3$ | eV<br>(h∙c)<br>(E3−E2)∙e\<br>6543•10 <sup>-10</sup> |  | m.  |                    |                  |
| $E2 := -3.4$ $\lambda = H_{3} := 3$ $\lambda = H_{3} := 3$ | eV<br>(h∙c)<br>(E3−E2)∙e\<br>6543•10 <sup>-10</sup> |  | m.  |                    |                  |
| $E2 := -3.4$ $\lambda = H_{3} := 3$ $\lambda = H_{3} := 3$ | eV<br>(h∙c)<br>(E3−E2)∙e\<br>6543•10 <sup>-10</sup> |  | Im.       Im.         Im.       I |                    |                  |
| $E2 := -3.4$ $\lambda = H_{3} := 3$ $\lambda = H_{3} := 3$ | eV<br>(h∙c)<br>(E3−E2)∙e\<br>6543•10 <sup>-10</sup> |  | m.  |                    |                  |

| Problem 3.4  | ∞ Infinity                       |
|--|----------------------------------|
| The first line of the Lyman series in the hydrogen spectrum has the wavelength 1200 Angstrom.                          |                                  |
| Calculate the wavelength of the second line.   | 3                                |
| Solution:  | n<br>n<br>n<br>2                 |
| Lyman series (ultraviolet region:  |                                  |
| $1/\text{lamda} = R((1/1^2) - (1/n^2))$<br>Lyman series starts at 1 for the first term, for the second term n = 2,3,4, | Lyman Series<br>A1<br>1200 Å     |
| So moving up to the next line is $n = 2$ , from 1 to $n=2$ .   |                                  |
| $1/lamda1 = R ((1/1) - (1/2^2))$<br>= R (1 - (1/4)<br>= R (3/4).   | Hydrogen spectrum                |
| Next to the line up from 2 is 3.   |                                  |
| $1/\text{lamda2} = R ((1/1) - (1/3^2))$<br>= R (1 - (1/9)<br>= R (8/9).  |                                  |
| If we divide (1/lamda 1) by (1/lamda2) it wo<br>lamda2) through proportioning.   | ould give us the value of (1/    |
| (1/lamda1) / (1/lamda 2) = lamda 2 / lamda   | 1. Correct.                      |
| We know the value of lamda1 equal 1200 Ar get the wavelength of lamda2.  | ngstrom we then multipy to it to |
| (1/lamda2) / (1/lamda1) R(3/4) / R(8/9)  |                                  |
| = 27/32.   |                                  |
| lamda2 = (27/32) x 1200 = 1012.5 Angstron  | n Ans.                           |

Exercise by: K S Bogha. Part II Structure of the Hydrogen Atom and The Bohr Model.



|  | evolving around a stationary nucleus.  |
|--|--|
| The energy required to excite the electro<br>47.2 eV.  | on from the second orbit to the third orbit is   |
| What is the atomic number of the atom.   |  |
|  |  |
| Solution:  | En = (-2.2x10^-18)(Z^2/n^2) eV   |
| 47.2eV   | this becomes $(2^{1}2^{1}1^{1}2^{2}) \in V$  |
|  | $En = (-13.6)(Z^2/n^2) eV$   |
| electron   |  |
|  | Then for frequency and wavelength o the  |
| nucleus  | rbit 3 radiation in the transition from n2 to n1:  |
| orbit 2  | En = -13.6(Z^2) ((1/ni^2) - (1/(nf^2)) eV  |
| orbit 1  | It is -13.6 in the expression, though it was   |
| $ \land \land \checkmark //$   | made positive 13.6 as per textbook. However  |
|  | the formulae/expression in textbook is -13.6.  |
|  | Z is the atomic number.  |
|  | Electron is transitioning from 2 to 3, moving  |
|  | up or outward, this requires absorbtion of energy. This is the 47.2eV absorbed i.e.  |
|  | excitation energy. Here we set it as negative.   |
|  |  |
|  |  |
|  |  |
| Discussion: The convention we use is -v  |  |
|  | e for absorbed. The original expression was for  |
| radiated energy where ni <nf (number="" 2<="" td=""><td></td></nf>   |  |
| radiated energy where ni <nf (number="" 2<br="">we reverse it to make nf be placed at ni<br/>correctly for the answer. This worked wi</nf>   | e for absorbed. The original expression was for<br><3), here its absorbed the opposite direction. So<br>ie at the front. So the resulting sign worked out<br>ith 47.2eV set as positive.   |
| radiated energy where ni <nf (number="" 2<br="">we reverse it to make nf be placed at ni<br/>correctly for the answer. This worked wi<br/>Since, as we set 47.2 eV negative it wor</nf>  | e for absorbed. The original expression was for <3), here its absorbed the opposite direction. So ie at the front. So the resulting sign worked out  |
| radiated energy where ni <nf (number="" 2<br="">we reverse it to make nf be placed at ni<br/>correctly for the answer. This worked wi<br/>Since, as we set 47.2 eV negative it wor<br/>also. You verify.</nf>  | e for absorbed. The original expression was for<br><3), here its absorbed the opposite direction. So<br>ie at the front. So the resulting sign worked out<br>ith 47.2eV set as positive.<br>ked out correctly with the original expression   |
| radiated energy where ni <nf (number="" 2<br="">we reverse it to make nf be placed at ni<br/>correctly for the answer. This worked wi<br/>Since, as we set 47.2 eV negative it wor</nf>  | e for absorbed. The original expression was for<br><3), here its absorbed the opposite direction. So<br>ie at the front. So the resulting sign worked out<br>ith 47.2eV set as positive.<br>ked out correctly with the original expression   |
| radiated energy where ni <nf (number="" 2<br="">we reverse it to make nf be placed at ni<br/>correctly for the answer. This worked wi<br/>Since, as we set 47.2 eV negative it wor<br/>also. You verify.</nf>  | e for absorbed. The original expression was for<br><3), here its absorbed the opposite direction. So<br>ie at the front. So the resulting sign worked out<br>ith 47.2eV set as positive.<br>ked out correctly with the original expression   |
| radiated energy where ni <nf (number="" 2<br="">we reverse it to make nf be placed at ni<br/>correctly for the answer. This worked wi<br/>Since, as we set 47.2 eV negative it wor<br/>also. You verify.</nf>  | e for absorbed. The original expression was for<br><3), here its absorbed the opposite direction. So<br>ie at the front. So the resulting sign worked out<br>ith 47.2eV set as positive.<br>ked out correctly with the original expression<br>eValternate expression<br>Alternate with nf = 3 placed in front.   |
| radiated energy where ni < nf (number 2<br>we reverse it to make nf be placed at ni<br>correctly for the answer. This worked wi<br>Since, as we set 47.2 eV negative it wor<br>also. You verify.<br>En = -13.6(Z^2) ((1/nf^2) - (1/(ni^2))<br>7.2 eV = -13.6 (Z^2) ( (1/ni^2) - (1/nf^2)   | e for absorbed. The original expression was for<br><3), here its absorbed the opposite direction. So<br>ie at the front. So the resulting sign worked out<br>ith 47.2eV set as positive.<br>ked out correctly with the original expression<br>eValternate expression<br>Alternate with nf = 3 placed in front.<br>^2) ) 47.2 eV = - 13.6 (Z^2) ( (1/nf^2) - (1/ni^   |
| radiated energy where ni < nf (number 2<br>we reverse it to make nf be placed at ni<br>correctly for the answer. This worked wi<br>Since, as we set 47.2 eV negative it wor<br>also. You verify.<br>En = -13.6(Z^2) ((1/nf^2) - (1/(ni^2))<br>7.2 eV = -13.6 (Z^2) ( (1/ni^2) - (1/nf^2)   | The for absorbed. The original expression was for<br><3), here its absorbed the opposite direction. So<br>ie at the front. So the resulting sign worked out<br>ith 47.2eV set as positive.<br>Where with the original expression<br>eValternate expression<br>Alternate with nf = 3 placed in front.<br>$(1/nf^2) - (1/ni^2)$  |
| radiated energy where ni < nf (number 2<br>we reverse it to make nf be placed at ni<br>correctly for the answer. This worked wi<br>Since, as we set 47.2 eV negative it wor<br>also. You verify.<br>En = -13.6(Z^2) ((1/nf^2) - (1/(ni^2))<br>7.2 eV = -13.6 (Z^2) ( (1/ni^2) - (1/nf^2)   | The for absorbed. The original expression was for<br><3), here its absorbed the opposite direction. So<br>ie at the front. So the resulting sign worked out<br>ith 47.2eV set as positive.<br>Where with the original expression<br>eValternate expression<br>Alternate with nf = 3 placed in front.<br>$(1/nf^2) - (1/ni^2)$  |
| radiated energy where ni < nf (number 2<br>we reverse it to make nf be placed at ni<br>correctly for the answer. This worked wi<br>Since, as we set 47.2 eV negative it wor<br>also. You verify.<br>En = -13.6(Z^2) ((1/nf^2) - (1/(ni^2))<br>7.2 eV = -13.6 (Z^2) ( (1/ni^2) - (1/nf^2)   | The for absorbed. The original expression was for<br><3), here its absorbed the opposite direction. So<br>ie at the front. So the resulting sign worked out<br>ith 47.2eV set as positive.<br>Ked out correctly with the original expression<br>eValternate expression<br>Alternate with nf = 3 placed in front.<br>^2) ) 47.2 eV = - 13.6 (Z^2) ((1/nf^2) - (1/ni^2))<br>89 Z <sup>2</sup> eV RHS_alt := -13.6 $\cdot \left( \left( \frac{1}{nf^2} \right) - \left( \frac{1}{ni^2} \right) \right)$ |
| radiated energy where ni < nf (number 2<br>we reverse it to make nf be placed at ni<br>correctly for the answer. This worked wi<br>Since, as we set 47.2 eV negative it wor<br>also. You verify.<br>En = -13.6(Z^2) ((1/nf^2) - (1/(ni^2))<br>7.2 eV = -13.6 (Z^2) ((1/ni^2) - (1/nf^2))<br>ni := 2 nf := 3<br>RHS := -13.6 $\cdot \left( \left( \frac{1}{ni^2} \right) - \left( \frac{1}{nf^2} \right) \right) = -1.88$   | The for absorbed. The original expression was for<br><3), here its absorbed the opposite direction. So<br>ie at the front. So the resulting sign worked out<br>ith 47.2eV set as positive.<br>Ked out correctly with the original expression<br>eValternate expression<br>Alternate with nf = 3 placed in front.<br>^2) ) 47.2 eV = - 13.6 (Z^2) ((1/nf^2) - (1/ni^2))<br>89 Z <sup>2</sup> eV RHS_alt := -13.6 $\cdot \left( \left( \frac{1}{nf^2} \right) - \left( \frac{1}{ni^2} \right) \right)$ |
| radiated energy where ni < nf (number 2<br>we reverse it to make nf be placed at ni<br>correctly for the answer. This worked wi<br>Since, as we set 47.2 eV negative it wor<br>also. You verify.<br>En = -13.6(Z^2) ((1/nf^2) - (1/(ni^2))<br>7.2 eV = -13.6 (Z^2) ((1/ni^2) - (1/nf^2))<br>ni := 2 nf := 3<br>RHS := -13.6 $\cdot \left( \left( \frac{1}{ni^2} \right) - \left( \frac{1}{nf^2} \right) \right) = -1.88$   | The for absorbed. The original expression was for<br><3), here its absorbed the opposite direction. So<br>ie at the front. So the resulting sign worked out<br>ith 47.2eV set as positive.<br>Ked out correctly with the original expression<br>eValternate expression<br>Alternate with nf = 3 placed in front.<br>^2) ) 47.2 eV = - 13.6 (Z^2) ((1/nf^2) - (1/ni^2))<br>89 Z <sup>2</sup> eV RHS_alt := -13.6 $\cdot \left( \left( \frac{1}{nf^2} \right) - \left( \frac{1}{ni^2} \right) \right)$ |
| radiated energy where ni < nf (number 2<br>we reverse it to make nf be placed at ni<br>correctly for the answer. This worked wi<br>Since, as we set 47.2 eV negative it wor<br>also. You verify.<br>En = -13.6(Z^2) ((1/nf^2) - (1/(ni^2))<br>7.2 eV = -13.6 (Z^2) ((1/ni^2) - (1/nf^2))<br>7.2 eV = -13.6 (Z^2) ((1/ni^2) - (1/nf^2)) | The for absorbed. The original expression was for<br><3), here its absorbed the opposite direction. So<br>ie at the front. So the resulting sign worked out<br>ith 47.2eV set as positive.<br>Where with the original expression<br>eValternate expression<br>Alternate with nf = 3 placed in front.<br>$(1/nf^2) - (1/ni^2)$  |
| radiated energy where ni < nf (number 2<br>we reverse it to make nf be placed at ni<br>correctly for the answer. This worked wi<br>Since, as we set 47.2 eV negative it wor<br>also. You verify.<br>En = -13.6(Z^2) ((1/nf^2) - (1/(ni^2))<br>7.2 eV = -13.6 (Z^2) ((1/ni^2) - (1/nf^2))<br>ni := 2 nf := 3<br>RHS := -13.6 $\cdot \left( \left( \frac{1}{ni^2} \right) - \left( \frac{1}{nf^2} \right) \right) = -1.88$   | The for absorbed. The original expression was for<br><3), here its absorbed the opposite direction. So<br>ie at the front. So the resulting sign worked out<br>ith 47.2eV set as positive.<br>Ked out correctly with the original expression<br>eValternate expression<br>Alternate with nf = 3 placed in front.<br>^2) ) 47.2 eV = - 13.6 (Z^2) ((1/nf^2) - (1/ni^2))<br>89 Z <sup>2</sup> eV RHS_alt := -13.6 $\cdot \left( \left( \frac{1}{nf^2} \right) - \left( \frac{1}{ni^2} \right) \right)$ |
| radiated energy where ni < nf (number 2<br>we reverse it to make nf be placed at ni<br>correctly for the answer. This worked wi<br>Since, as we set 47.2 eV negative it wor<br>also. You verify.<br>En = -13.6(Z^2) ((1/nf^2) - (1/(ni^2))<br>7.2 eV = -13.6 (Z^2) ((1/ni^2) - (1/nf^2))<br>7.2 eV = -13.6 (Z^2) ((1/ni^2) - (1/nf^2)) | The for absorbed. The original expression was for<br><3), here its absorbed the opposite direction. So<br>ie at the front. So the resulting sign worked out<br>ith 47.2eV set as positive.<br>Ked out correctly with the original expression<br>eValternate expression<br>Alternate with nf = 3 placed in front.<br>^2) ) 47.2 eV = - 13.6 (Z^2) ((1/nf^2) - (1/ni^2))<br>RHS_alt = -13.6 (((1/nf^2) - (1/ni^2)))<br>RHS_alt = 1.889<br>Z_squared <sub>alt</sub> := $\frac{47.2}{1.8889}$ = 24.988   |

| Which state of the           | triply ionized Perullium has the same arbital radius as that of the  |
|------------------------------|--|
| ground state of the          | triply ionized Beryllium has the same orbital radius as that of the drogen?  |
| Compare the energ            | gies of the two states.  |
| Solution:                    |  |
| Triply ionised Bery          | llium: Be+++.  |
|                              | Gaining 3 positive ions.<br>Z = 4.   |
|                              | n?   |
| Hydrogen H: $Z = 1$<br>n = 1 |  |
|                              | radius of the nth bohr orbit<br>) ((h/2pi)^2)n^2 ) / (Z(e^2)m)   |
|                              | antum number, substituting epsilion_0, (h/2pi), e, and m for their pression rn results in  |
| rn = 0.53 (n^2/Z)            | Angstrom   |
| have their radius e          | of Be+++ that has the same orbital radius as H would<br>qual. The radius is dependent on Z. Since we are<br>hilar equations the constant term 0.53 is cancelled.                 |
| ((n^2)/Z) Be+++              | = ((n^2)/Z) H  |
| ((n^2)/4) = ((1^2            | 2)/1) = 1  |
| ((n^2) = 4                   |  |
| n = 2.                       |  |
| n Be+++ = 2 (2nd             | l orbit) Ans.  |
|                              | nergies of the 2 states by having one divide the other, or ydrogen goes into the energy of Be+++.  |
| En proportional to:          | $ \begin{array}{c} -13.6(Z^{2}/n^{2})in \ comparison \ -13.6 \ cancelled \ out. \\ (Z^{2}/n^{2}) \\ = \ ((4^{2})/(2^{2})) \ / \ ((1^{2})/(1^{2})) \\ = \ 4 \ / \ 1 \end{array} $ |

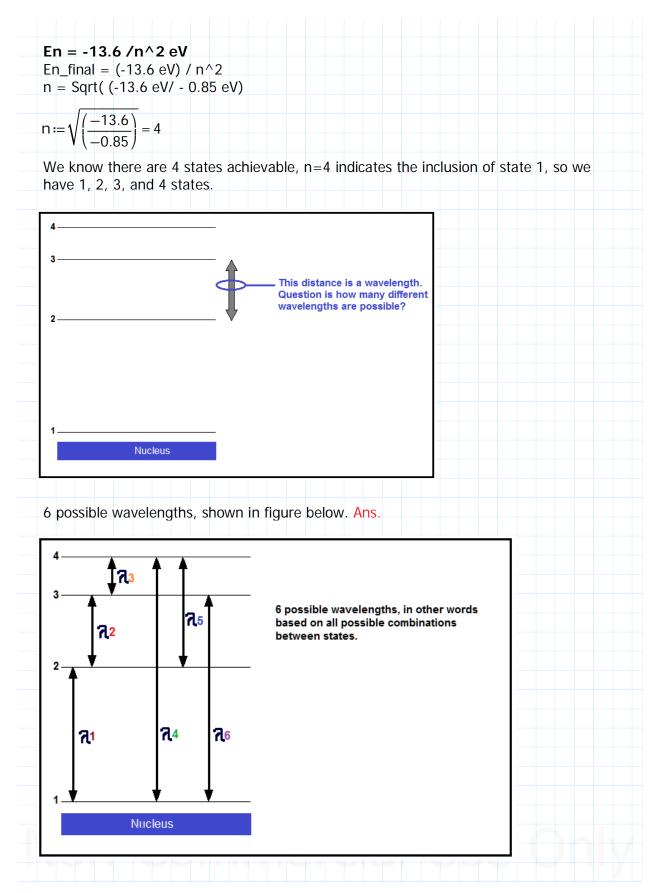
| Which state of the do<br>of hydrogen state en<br>Hint: Compare the o   | ergy of the Hydrog        |                 | e energy as the gr  | round state      |
|--|---------------------------|-----------------|---------------------|------------------|
| Hint: Compare the o  |                           |                 |                     |                  |
|  | rbital radii of the tw    | wo states.      |                     |                  |
| Solution:  |                           |                 |                     |                  |
| Doubly ionised Lithiu  | m: Li++.<br>Gaining 2 pos | itivo ions      |                     |                  |
|  | Z = 3.                    |                 |                     |                  |
| Hydrogen H: $Z = 1$ .<br>n = 1.  | n?                        |                 |                     |                  |
| rn = 0.53 (n^2/Z) A  | ngstrom                   |                 |                     |                  |
| rn = (n^2/Z) Angstr  | om. When compari          | ing two element | s the constant 0.5  | 53 cancelled out |
| $((n^2)/3)$ Li++ = ((n<br>Lets try n = 3<br>$((3^2)/3)$ = ((<br>3 = 1. | 1^2)/1) = 1               |                 |                     |                  |
| When $n = 3$ for Li, n<br>The $n = 3$ state of Li                      |                           |                 | ne n = 1 state of F | 4.               |
|  |                           |                 |                     |                  |
|  |                           |                 |                     |                  |
|  |                           |                 |                     |                  |
|  |                           |                 |                     |                  |
|  |                           |                 |                     |                  |
|  |                           |                 |                     |                  |
|  |                           |                 |                     |                  |
|  |                           |                 |                     |                  |
|  |                           |                 |                     |                  |
|  |                           |                 |                     |                  |

Problem 3.9 Hydrogen atom in its ground state is excited by means of a monochromatic radiation of wavelength 970.6 Angstrom. How many different wavelengths are possible in the resulting emission spectrum? Find the longest wavelength amongst these. Solution: ? <--Results in how many states Monochromatic radiation of nucleus wavelength 970.6 Angstrom. orbit 1 n=1 is the ground state Nucleus Energy of the radiation quantum (hv): Since we are given the wavelength of the excitor source we have to use lamda expression for hv, hc/lamda to calculate E.  $\lambda = 970.6 \cdot 10^{-10}$ m^-1  $h = 6.6 \cdot 10^{-34}$  Js  $\mathsf{E} := \frac{(\mathsf{h} \cdot \mathsf{c})}{\lambda} = 2.04 \cdot 10^{-18}$ J  $E_eV := \frac{E}{eV} = 12.75$  eV energy of the monochromatic radiation.  $E_Hydrogen := -13.6$ eV in the unexcited state at n = 1. Final or resulting energy after the excitation energy is struck on the Hydrogen atom:  $En_{final} = E_Hydrogen + E_eV = -0.85 eV$ **Remember:** This formula  $En = -13.6 / n^2 eV$  for Hydrogen at state n = 1 can produce the spread of lines, or states. The variable 'n' should solve for the core of our

problem. This is the spread of states the Hydrogen atom generates and each state has its wave length. *Usually comes to realisation as we solve the difficulty!* 

Quantum Mechanics A Textbook for Undergraduates. Mahesh C. Jain. PHI. Department of Physics, Hindu College University of Delhi. To Support Studies In: Modern Physics by S.N. Ghosal.

Purpose: <u>Quantum Mechanics for 'Power Plant Engineering' Studies.</u> Exercise by: K S Bogha. Part II <u>Structure of the Hydrogen Atom and The Bohr Model.</u>



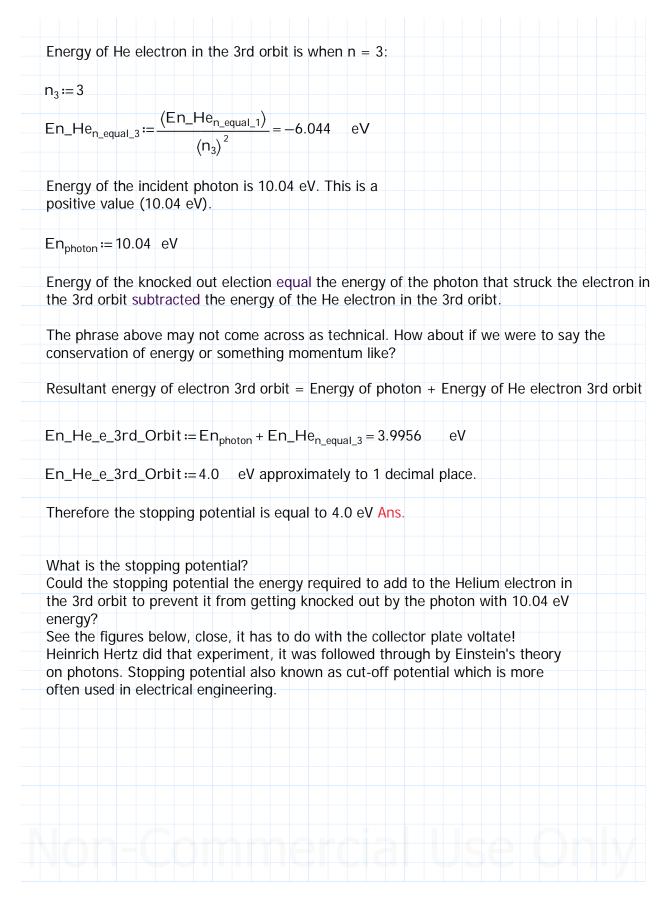
| Returning to the expression   |  |
|---|--|
| hc/lamda = E = En   |  |
| lamda = hc/En   |  |
| En is the value at the line n, the wa   | avelength is across 2 energy state lines.  |
|   | ifficulty is the value of energy across 2 states   |
| corresponding to the wavelength.  | 5  |
|   |  |
| En4 - En3 for example would be lar  |  |
|   | = (hc)/(En4 - En3). Correct. Just so happens its   |
| the smallest energy difference.   |  |
| The value of lamda gots largor w  | vith a smaller energy difference (En+1 - En).  |
| Correct Again!  |  |
|   |  |
| So if we know the smallest energy   | difference of the 6 possible energy differences, ther  |
|   | value into the expression, solving for lamda we  |
| should get the longest wavelength   |  |
|   |  |
| Marken and the line of the second second  |  |
| We know from the Hydroden energy  | iv states values given in textbooks, here  |
|   | y states values given in textbooks, here   |
| provided on page 1, in this case it i   | is between En4 and En3.  |
| provided on page 1, in this case it i<br>However we can also easily calcula   | is between En4 and En3.  |
| provided on page 1, in this case it i   | is between En4 and En3.  |
| provided on page 1, in this case it i<br>However we can also easily calculate<br>$En = -13.6eV/n^2$ .   | is between En4 and En3.  |
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Quantum Mechanics A Textbook for Undergraduates. Mahesh C. Jain. PHI. Department of Physics, Hindu College University of Delhi. To Support Studies In: Modern Physics by S.N. Ghosal. Purpose: Quantum Mechanics for 'Power Plant Engineering' Studies. Exercise by: K S Bogha. Part II Structure of the Hydrogen Atom and The Bohr Model. Problem 3.10 In a singly-ionized Helium atom the electron is in the third orbit. A photon of energy 10.04 eV knocks out the electron. Calculate the stopping potential of the electron. The ionization energy of Hydrogen atom is 13.6 eV. Solution: Knocks-out electron Discussion: Knocks-out electron? Current is a flow of electron. We have photon knocking out 1 electron in Photon of energy 10.04eV the 3rd orbit of an ionised Hydrogen atom. orbit 3 Is this not an introduction to Solar PV cells? orbit 2 I say it is. It takes a lot of photons and electrons to get an appreciable electric current, and that is found in many solar (PV)panels. Atom: He Z: 2 Ionization: 1, atom is positively charged He+ (say nucleus is +pos charged). Electron: 3rd orbit is found 1 electron (electron is negatively charged) Helium ionization energy is found based on comparision to Hydrogen atom ionisation energy.  $E(He) = (Z^2) HeV$ Z<sub>He</sub>:=2 H<sub>ionization energy</sub>:=13.6 eV  $He_{ionization\_energy} := H_{ionization\_energy} \cdot (Z_{He}^{2}) = 54.4$ eV at ground state (for He+) Note: n = 1 ground state is first orbit n = 2 is second orbit n = 3 is third orbit n = 4 is the fourth orbit and so on .... Energy of He electron in the 1st orbit is when n = 1: He\_n equal 1 =  $-(He+)/n^2$  $n_1 := 1$  $En_He_{n_equal_1} := \frac{-(He_{ionization_energy})}{(n_1)^2} = -54.4 \text{ eV}$ 

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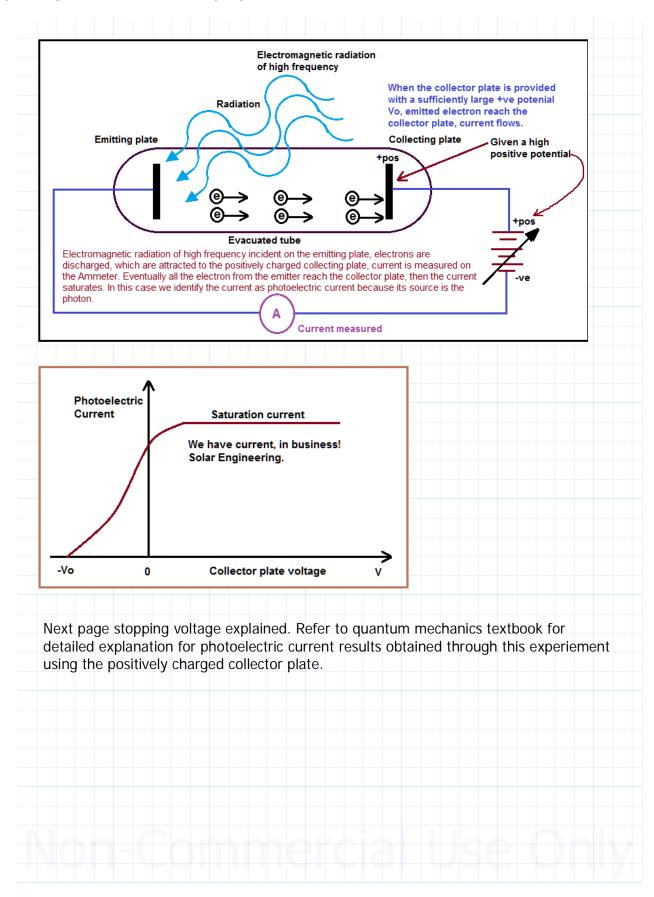
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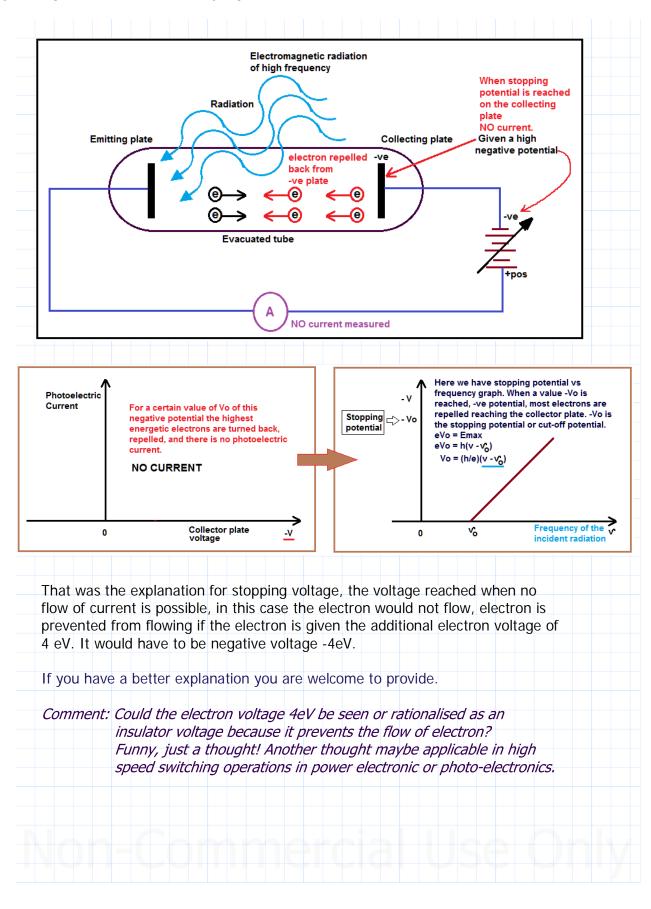
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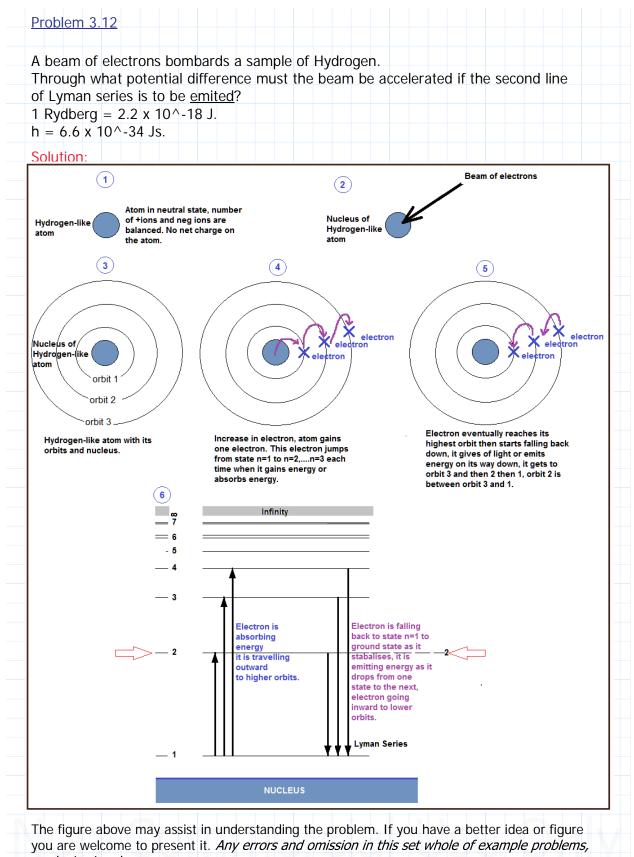
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|  | e to the ground state.  |
|--|---|
| 1 Rydberg<br>h = 6.6 x 1                     | = 2.2 x 10^-18 J.<br>0^-34 Js.  |
| Solution:                                    |   |
| R≔2.2•10                                     | -18 J   |
| H-like atom                                  | ionisation energy = 4 x Rydberg.  |
| E <sub>H_like</sub> ≔4∙                      | $R = 8.8 \cdot 10^{-18}$ J  |
| Energy of ti<br>negative sig                 | he electron in the ground state (n=1) is equal to E_H_like but with a gn.                                   |
| E1 <sub>H_like</sub> ≔–                      | $E_{H_{Like}} = -8.8 \cdot 10^{-18}$ J  |
| Energy of tl                                 | he electron in the 2nd orbit (n=2) is equal to:   |
| n2:=2  |   |
| E2 <sub>H_like</sub> ≔_                      | $\frac{-E_{H_{-}like}}{n2^{2}} = -2.2 \cdot 10^{-18} \qquad J$  |
|  | previous example problem using the energy between 2 states to e wavelength. Radiation emitted is (E2 - E1). |
| (hc)/lamda                                   | = E2 - E1, therefore lamda (wavelength) = (hc)/(E2 - E1).   |
| $\lambda \coloneqq \frac{1}{(E2_{H_{III}})}$ | $\frac{(h \cdot c)}{k_e - E 1_{H_{-}IIke}} = 3 \cdot 10^{-8}$ m^-1.   |
| $\lambda := 300 \cdot 10^{\circ}$            | <sup>-10</sup> Angstrom Ans.  |



apologies in advance.

|                               | series of the Hydrogen atom starts with the orbit $n=1$ .  |
|-------------------------------|--|
| Energy                        | is absorbed on the when going from $n=1$ to higher orbits.<br>is emitted when going from higher orbits to $n=1$ ie to the ground state or<br>state. This problem statement is concerned about energy emitted.  |
| It gains<br>reaches           | rt with bombarding an atom that is not ionised. Lets say in its neutral state.<br>s energy and starts sending the electron outward, energy absorbed. When it<br>s it highest state it then falls back to ground state. The Lyman series, has $n = 2$<br>on $n=3$ and $n=1$ .   |
|                               | nd the potenial difference between state $n=3$ and $n=1$ this would provide us ergy that is <u>bound between <math>n=3</math> and <math>n=1</math></u> .   |
|                               | oblem statement ask what must the 'beam' be acceleratedthere is only one hat which bombards the sample of Hydrogen.  |
| gaining<br>bound,<br>differer | y acceleration occurs when the electron progress from lower to higher states,<br>energy from the beam the electron accelerates. So if we calculate the energy<br>eV, betwen states 3 and 1 that may provide us the minimum potential<br>nce. We are NOT computing acceleration of electrons, rate of change of speed/<br>y of electrons, rather just the potential difference. Little simpler! |
| revious<br>omments            | s: You verify on the sign convention.  |
|                               | Its not exactly like Kirchchoff's Law in electric circuit's convention.Lets say if we set the convention, radiated outward is +ve sign,then absorbed would be -ve sign.Provided the ni and nf line up with theorder of energy direction movement.  |
| Binding                       | energy of the atom is n=1 state:   |
| En1≔                          | 13.6 eV.   |
| Binding                       | energy of the atom in n=3 state:   |
| En3≔-                         | $\frac{\text{En1}}{(3^2)} = 1.511 \text{ eV.}$   |
| Energy                        | required to leap or jump from n=1 to n=3 state:  |
|                               | f_n1_n3:= En1 – En3 = 12.089 eV. Ans.  |

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ie electron (m) + nucleus (M), it would be  $1 \times 1 = 1$ .

If the electron was a percentage smaller than nucleus, then we could say m x M. where m is a fraction of M, example  $0.0001 \times 1.0 = 0.0001$  results with the same mass of electon. It is the mass of electron (numerator), to 'nucleus and electron' in denominator we seek. This expression provides a fix or remedy.

v (frequency) = 1/(lamda) = R(infinity) (Z^2) ((1/n1^2) - (1/n2^2))

 $R(infinity) = (m/(4 \text{ pi } c (h'^3))) (e^2/(4 \text{ pi epsilon}_o)^2.$  h' = h/2pi. $R(infinity) \text{ is the Rydberg contant} = 1.09737 \times 10^{-7} \text{ m}^{-1}.$ 

Now correcting for the Rydberg constant for Hydrogen

 $R_H = R(infinity) / (1 + (m/M_H) M_H mass of Hydrogen nucleus.$ 

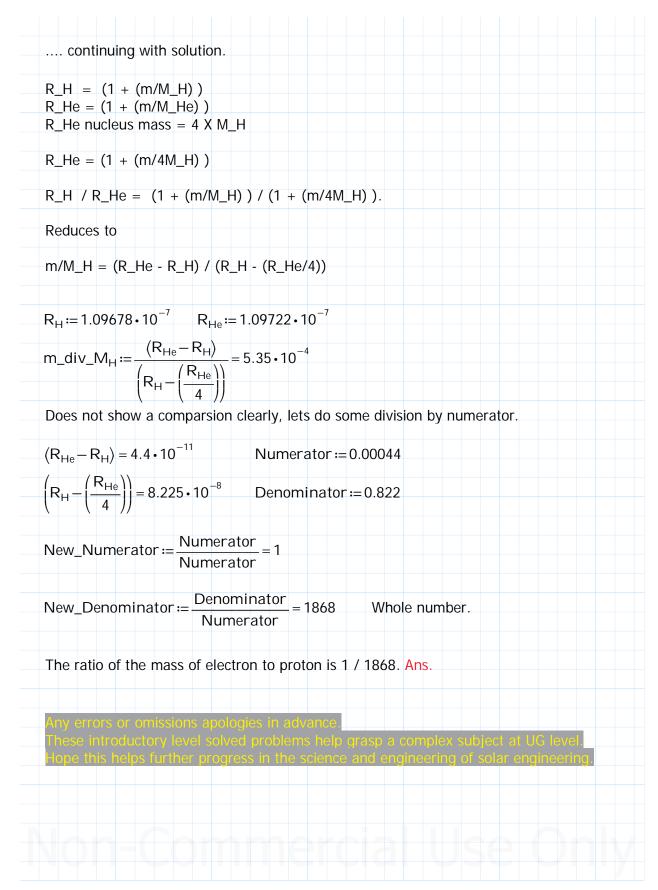
Now  $R_H = 1.09678 \times 10^{-7}$  is closer to an empirical value 1.096776 x 10<sup>-7</sup> that it is closer to R(infinity) 1.09737 x 10<sup>-7</sup>.

Look up in modern physics or quantum mechanics textbook on the full notes on this

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Exercise by: K S Bogha. Part II Structure of the Hydrogen Atom and The Bohr Model.



Textbook: Modern (Atomic) Physics Vol I. Author:S.N. Ghosal Publisher: S. Chand. Chapter 9: Wave particle duality; Hisenberg's uncertainty principle. 9.1 Particle and waves 9.2 Phase and group velocities 9.3 Particle wave 9.4 Relation between phase and group velocity of de Broglie waves 9.5 Discovery of matter waves: Davisson and Germer's experiment 9.6 G.P. Thompson's experiment 9.7 Effect of refraction of the electron beam 9.8. De Broglie wavelength of high energy electrons 9.9 Electron microscope 9.10 Need for a new mechanics for the sub-atomic particles 9.11 Particles and wave packets 9.12 Nature of matter waves 9.13 Uncertainty relations 9.14 Gamma-ray microscope experiment 9.15 Applicability of classical and quantum concepts 9.16 Principle of superposition The solved example problems in the textbook 'Quantum Mechanics: A textbook for Undergraduates' by Mahesh C. Jain are used for UG subject matter comprehension. April 8, 1999 • Phys. Rev. Focus 3, 21 Focus: The de Broglie Wavelength of a Packet of Light Texas Christian University Double Slit. Unlike this typical interference experiment where plane waves fall on a double slit, physicists have been able to observe the interference pattern from a wave packet consisting of just two photons, measuring the effective wavelength in the process. Show Less Source: April 8, 1999. Phy. Rev. Focus 3,21.

Example problems here are primarily de Broglies Hypothesis related. *Comments: Chapter 9, Ghosal textbook, contents are excellent in understanding the background for therory, its mathematics, it includes derivations, figures, and explanations. You may find this in similar textbooks today.* 

| Constants:  |  |
|---|--|
| $h := 6.63 \cdot 10^{-34}$  | Js   |
| $c := 3 \cdot 10^8$<br>$eV := 1.6 \cdot 10^{-19}$                                   | m/s  |
| $eV := 1.6 \cdot 10^{17}$   | J  |
| $e := 1.6 \cdot 10^{-19}$   | J  |
| $m_{neutron} = 1.675 \cdot 10$  | kg   |
| $m_{neutron} \coloneqq 1.675 \cdot 10$ $m_{electron} \coloneqq (9.1 \cdot 10^{-7})$ | <sup>31</sup> ) kg   |
| Problem 4.1   |  |
| Find the de Broglie of 100V.  | wavelength of electrons accelerated through a potential difference       |
| Solution:   |  |
| Wavelength lamda =<br>OR  | = h / Sqrt(2 m e V)  |
|   | = (h / Sqrt(2 m e)) (1/(Sqrt(V)) Angstrom<br>= 12.3 (1/Sqrt(V)) Angstrom |
| V := 100 Volt   |  |
| $\lambda \coloneqq \frac{12.3}{\sqrt{V}}  \text{Angstron}$                          | m  |
| $\lambda = 1.23$ Angstro  | m Ans.   |
| Problem 4.2   |  |
| Find the de Broglie   | wavelength of electrons moving with a kinetic energy of 100 eV.          |
| Solution:   |  |
| _   | = h / (Sqrt( 2 m K))   |
| $\lambda := $ (h)   | $= 1.229 \cdot 10^{-10}$   |
| $\sqrt{(2 \cdot m_{electron} \cdot 1)}$   | $\frac{1}{100 \cdot eV} = 1.229 \cdot 10^{-10}$                          |
| $\lambda \coloneqq 1.229 \ Angstrom$  | n Ans.   |
|   |  |
|   |  |

| Wł         | nat should be the kinetic energy of a neutron in eV so that its associated de Broglie  |
|------------|--|
| wa         | velength is 1.4 x 10^-10 m?  |
| Ма         | ss of neutron = 1.675 x 10^-27 kg.   |
| So         | lution:  |
|            | avelength lamda = h / (Sqrt( 2 m K)) rearrainging for K (kinetic energy)<br>KE or K = h^2 / (2 m lamda^2)  |
|            | = 1.4 • 10 <sup>-10</sup> m  |
| _          | $(h^2)$  |
| KE         | $:= \frac{(h^2)}{2 \cdot m_{\text{neutron}} \cdot \lambda^2}$  |
|            | $L = 6.695 \cdot 10^{-21}$ J   |
| KE         | $E := \frac{KE}{eV} = 0.0418$ eV   |
|            |  |
| KE         | $E := 4.18 \cdot 10^{-2}$ eV Ans.  |
| <u>Pro</u> | blem 4.4   |
| An         | electron in a hydrogen-like atom, is in an excited state.  |
|            | nas a total energy of -3.4 eV.   |
| Са         | culate (a) the kinetic energy  |
|            | (b) the de Broglie wavelength of the electron.   |
| So         | lution:  |
|            | Hydrogen like atom,similar to the energy level diagram of Hydrogen.  |
|            | = 1, En1-H = - 13.6 eV -ve sign indicating energy absorbed travelling downward   |
|            | what was reffered in part II as energy direction (movement)  |
|            | the usual direction is downward, absorbing energy.   |
| He         | re this H-like atom has a total value of -3.4 eV.  |
|            | is -3.4eV may be considred in the positive sign if its seen like this;   |
| KE         | or K = - (total energy of H-like electron E)<br>= - (-3.4 eV)  |
|            | = 3.4 eV Ans.  |
|            | <b>mment:</b> You may ask what is the purpose of this exercise? It may be just to show how to<br>erpret the KE's direction another way. You may say why should be KE be of an opposite sign to |
|            | total energy given? Its in an excited state, it was given excitation, its initial energy may been  |
| +v         | e, the excitation created a change it got -ve energy, finally we interpter the final energy as   |
|            | sitive. Something to consider before labelling a final answer to any of QMs problems.Ok, now   |
| loc        | k what happens in (b) part of solution, if it were negative -3.4 the answer would be a complex   |

| wavelengt  | h lamda = h / (Sqrt( 2 m K))  |
|--|---|
| K≔3.4 eV   | note the postive sing used here, otherwise sqrt of a -ve number returns a complex number. Unit eV included in expression. |
| <u> ۲.–</u>  | (h) $-6.663 \cdot 10^{-10}$   |
| $\sqrt{2 \cdot r}$   | $\frac{(h)}{n_{\text{electron}} \cdot K} = 6.663 \cdot 10^{-10}$  |
| $\lambda \coloneqq 6.663$                                      | Angstrom Ans.   |
| Problem 4.5  |   |
| Find the kinet<br>Angstrom.                                    | ic energy of a neutron in electron-volt if its de Broglie wavelength is 1.0   |
| Mass of neutr  | $ron = 1.674 \text{ x } 10^{-27} \text{ kg}; h = 6.60 \text{ x } 10^{-34} \text{ Js}.$                                    |
| Solution:<br>m <sub>neutron</sub> ==                           | 1.674 • 10 <sup>-27</sup> kg  |
| $h \coloneqq 6.60 \cdot 1$ $\lambda \coloneqq 1 \cdot 10^{-1}$ | 0 <sup>-34</sup> Js<br><sup>0</sup> Angstrom  |
|  | $=\frac{\left(\frac{(h^2)}{2 \cdot m_{\text{neutron}} \cdot \lambda^2}\right)}{eV} = 0.081  eV$                           |
|  | $=\frac{\left(2\cdot m_{\text{neutron}}\cdot\lambda^{2}\right)}{eV}=0.081 \qquad eV$                                      |
| KE <sub>neutron</sub> ∺  |   |
|  | $=8.1 \cdot 10^{-2}$ eV Ans.  |
|  |   |
|  |   |
|  |   |
|  |   |
|  |   |

| A ball of  | f mass 1  | o a is   | s mov   | /ina v  | with   | a spe   | ed of  | 1 m/s   |  |   |  |   |  |           |       |   |
|--|---|--|---|---|--|---|--|---|--|---|--|---|--|-----------|-------|---|
|  | te the de   | 0  |   | 0   |  |   |  |   |  |   |  |   |  |           |       |   |
|  | effect o  |  |   |   | 0  |   |  |   |  | tally   | >  |   |  |           |       |   |
|  |   |  |   |   |  |   |  |   |  |   |  |   |  |           |       |   |
| Solution   | 1:  |  |   |   |  |   |  |   |  |   |  |   |  |           |       |   |
|  |   |  |   |   |  |   |  |   |  |   |  |   |  |           |       |   |
| Waveler  | ngth lam  | da =   | h/(   | ímv)  | ref  | er to   | mode   | rn phy  | /sics t  | textb   | ook  | or an   | n tex  | tbo       | ok fo | r |
|  | 5   |  |   |   |  |   | nula/e   |   |  |   |  |   |  |           |       |   |
|  |   |  |   |   |  |   |  |   |  |   |  |   |  |           |       |   |
| m_ball:  | = 10 • 10 <sup>-</sup>  | -3   | kg  | 10  | g =  | 10/1  | 000 g  | = 0.0   | 1 = 1  | 0 *1  | 0^-3   |   |  |           |       |   |
|  |   |  | Ū   |   | Ū  |   | Ū  |   |  |   |  |   |  |           |       |   |
| v≔1.0  | m/s   |  |   |   |  |   |  |   |  |   |  |   |  |           |       |   |
|  |   |  |   |   |  |   |  |   |  |   |  |   |  |           |       |   |
| λ:=  | h<br>_ball•v)   | = 6.6  | •10   | -32   | m  | Ans.  |  |   |  |   |  |   |  |           |       |   |
| (m_  | _ball•v)  |  |   |   |  |   |  |   |  |   |  |   |  |           |       |   |
|  |   |  |   |   |  |   |  |   |  |   |  |   |  |           |       |   |
|  |   |  |   |   |  |   |  |   |  |   | ry cr  | ما الد  | nat  | h         |       |   |
| The de   | Broalie v   | vavel  | enath   | n was   | s calo   | culate  | ed abo   | ve, its   | a ve   | v ve  | IV SI  |   | JUNGL  |           |       |   |
|  | Broglie v<br>/isible. So  |  | -   |   |  |   |  |   |  | -   | -  |   | -  |           |       |   |
| its not v  | isible. So  | o if w   | ie we   | ere to  | con  | duct  | an exp   | periem  | ent o  | n thi   | s bal  | l to d  | irec   | tly       |       |   |
| its not v<br>see effe  | -   | o if w<br>s wav  | ve we<br>veleng   | ere to<br>gth w   | con<br>ould  | duct a<br>be u  | an exp<br>inreali:   | oeriem<br>stic. 1   | ent o<br>⁻he w   | n thi<br>avel   | s bal<br>enth  | l to d<br>is so   | irect<br>muc   | tly       |       |   |
| its not v<br>see effe<br>smaller   | isible. So  | o if w<br>s wav<br>ball  | ve we<br>veleng<br>of ma  | ere to<br>gth w<br>ass 1  | con<br>ould  | duct a<br>be u  | an exp<br>inreali:   | oeriem<br>stic. 1   | ent o<br>⁻he w   | n thi<br>avel   | s bal<br>enth  | l to d<br>is so   | irect<br>muc   | tly       |       |   |
| its not v<br>see effe<br>smaller<br>observe  | visible. So<br>ect of this<br>than the<br>ed experin  | o if w<br>s wav<br>ball<br>ment  | ve we<br>veleng<br>of ma<br>ally.   | ere to<br>gth w<br>ass 1<br>Ans.  | con<br>/ould<br>0 gra  | duct<br>be u<br>am/s  | an exp<br>inreali:<br>dimer  | beriem<br>stic. T<br>Isions.  | ent o<br><sup>-</sup> he w<br>The  | n thi<br>avel<br>effec  | s bal<br>enth<br>ct car  | l to d<br>is so<br>nnot l   | irect<br>muc<br>be   | tly<br>ch |       |   |
| its not v<br>see effe<br>smaller<br>observe  | visible. So<br>ect of this<br>than the<br>ed experin<br><b>DR Comm</b>  | o if way<br>s way<br>ball<br>ment  | ve we<br>veleng<br>of ma<br>ally.<br><b>OR J</b>  | ere to<br>gth w<br>ass 1<br>Ans.  | con<br>/ould<br>0 gra<br><i>If thi</i> :                                 | duct<br>be u<br>am/s<br>s <i>ball</i>   | an exp<br>inreali<br>dimer<br><i>was ro</i>  | beriem<br>stic. T<br>isions.  | ent o<br>⁻he w<br>The<br>⁺ <i>a cor</i>  | n thi<br>avel<br>effec<br>nstan   | s bal<br>enth<br>ct car<br><i>t spe</i> e  | I to d<br>is so<br>nnot I<br>ed on  | irect<br>muc<br>be<br><i>a fla</i>                                     | tly<br>ch |       |   |
| its not v<br>see effe<br>smaller<br>observe<br><i>Input C</i><br><i>table, co</i>  | visible. So<br>ect of this<br>than the<br>ed experim<br><b>DR Comm</b><br>puld a high   | o if w<br>s wav<br>ball<br>ment<br><b>nent</b>                                 | ve we<br>veleng<br>of ma<br>ally.<br><b>OR J</b><br>pargeo  | ere to<br>gth w<br>ass 1<br>Ans.<br>oke:  | con<br>vould<br>0 gra<br>If this<br>tron                                 | duct<br>be u<br>am/s<br>s ball<br>beam  | an exp<br>inreali<br>dimer<br><i>was ro</i><br><i>aimed</i>  | beriem<br>stic. T<br>isions.<br><i>Illing at</i><br>at it a   | ent o<br>The w<br>The<br><i>a cor</i><br><i>t the s</i>  | n thi<br>avel<br>effec<br>nstan   | s bal<br>enth<br>ct car<br><i>t spee</i>   | I to d<br>is so<br>not I<br>ed on<br>d of 1   | irect<br>muc<br>be<br><i>a fla</i>                                     | tly<br>ch |       |   |
| its not v<br>see effe<br>smaller<br>observe<br><i>Input C</i><br><i>table, corrunning</i>  | visible. So<br>ect of this<br>than the<br>ed experim<br><b>DR Comm</b><br><i>puld a high</i><br><i>along side</i>                           | o if w<br>s wav<br>ball<br>ment<br>hly ch<br>c, rev                            | ve we<br>velenç<br>of ma<br>ally.<br><b>OR J</b><br>pargec<br>veal al   | ere to<br>gth w<br>ass 1<br>Ans.<br><i>oke:</i><br>d election<br>ny pai                   | con<br>ould<br>0 gra<br><i>If this</i><br><i>tron i</i>                  | duct a<br>be u<br>am/s<br><i>s ball</i><br>beam<br>on a                           | an exp<br>inreali<br>dimer<br><i>was ro</i><br><i>aimed</i><br>photog  | beriem<br>stic. T<br>sions.<br><i>Iling at</i><br><i>at it a</i><br><i>raphic</i>   | ent o<br>The w<br>The<br><i>a cort</i><br>t the s<br>plate   | n thi<br>avel<br>effec<br>nstan<br>same<br>? Like                             | s bal<br>enth<br>ct car<br><i>t spee</i><br><i>spee</i><br><i>e in th</i>                  | l to d<br>is so<br>nnot l<br>ed on<br>d of 1<br>ee  | irect<br>muc<br>be<br><i>a fla</i><br><i>m/s</i>                       | tly<br>ch |       |   |
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| its not v<br>see effe<br>smaller<br>observe<br><i>Input C</i><br><i>table, cc</i><br><i>running</i><br><i>Thompse</i><br><i>electron</i><br><i>set it up</i> | visible. So<br>ect of this<br>than the<br>ed experim<br><b>DR Comm</b><br>build a high<br>along side<br>on experin<br>flow thro<br>moving v | o if w<br>s wav<br>ball<br>ment<br>hly ch<br>e, rev<br>emen<br>ugh i<br>with a | ve we<br>veleng<br>of ma<br>ally.<br><i>OR Ja</i><br><i>pargeo</i><br><i>real al</i><br><i>t. The</i><br><i>t. May</i><br><i>spee</i> | ere to<br>gth w<br>ass 1<br>Ans.<br>oke:<br>d elec<br>ny pa<br>e ball<br>vbe a<br>ed of 1 | con<br>ould<br>ogra<br>If this<br>tron<br>ttern<br>be m<br>patte<br>1m/s | duct<br>be u<br>am/s<br><i>s ball</i><br>beam<br>on a<br>ade o<br>ern on<br>may l | an exp<br>inrealis<br>dimer<br><i>was ro</i><br><i>aimed</i><br><i>photog</i><br><i>of a ma</i><br><i>o the pl</i><br><i>be a hu</i> | beriem<br>stic. T<br>isions.<br><i>Illing at</i><br><i>at it a</i><br><i>raphic</i><br><i>terial t</i><br><i>late ma</i><br><i>irdle.</i> T | ent o<br>The w<br>The<br><i>a cor</i><br><i>t the s</i><br><i>plate</i><br><i>plate</i><br><i>type th</i><br><i>tyr reve</i> | n thi<br>avel<br>effec<br>nstan<br>same<br>? Like<br>nat w<br>eal so<br>ay no | s bal<br>enth<br>ct car<br>spee<br>e in th<br>ould p<br>ometh<br>ot wo                     | l to d<br>is so<br>nnot l<br><i>ed of 1</i><br><i>he</i><br><i>berming. E</i><br><i>rk if b</i> | irect<br>muc<br>be<br><i>a fla</i><br><i>m/s</i><br>t<br>But to<br>oth | tly<br>ch |       |   |
| its not v<br>see effe<br>smaller<br>observe<br><i>Input C</i><br><i>table, cc</i><br><i>running</i><br><i>Thompse</i><br><i>electron</i><br><i>set it up</i> | visible. So<br>ect of this<br>than the<br>ed experin<br><b>DR Comm</b><br>build a higu<br>along side<br>on experin<br>flow thro             | o if w<br>s wav<br>ball<br>ment<br>hly ch<br>e, rev<br>emen<br>ugh i<br>with a | ve we<br>veleng<br>of ma<br>ally.<br><i>OR Ja</i><br><i>pargeo</i><br><i>real al</i><br><i>t. The</i><br><i>t. May</i><br><i>spee</i> | ere to<br>gth w<br>ass 1<br>Ans.<br>oke:<br>d elec<br>ny pa<br>e ball<br>vbe a<br>ed of 1 | con<br>ould<br>ogra<br>If this<br>tron<br>ttern<br>be m<br>patte<br>1m/s | duct<br>be u<br>am/s<br><i>s ball</i><br>beam<br>on a<br>ade o<br>ern on<br>may l | an exp<br>inrealis<br>dimer<br><i>was ro</i><br><i>aimed</i><br><i>photog</i><br><i>of a ma</i><br><i>o the pl</i><br><i>be a hu</i> | beriem<br>stic. T<br>isions.<br><i>Illing at</i><br><i>at it a</i><br><i>raphic</i><br><i>terial t</i><br><i>late ma</i><br><i>irdle.</i> T | ent o<br>The w<br>The<br><i>a cor</i><br><i>t the s</i><br><i>plate</i><br><i>plate</i><br><i>type th</i><br><i>tyr reve</i> | n thi<br>avel<br>effec<br>nstan<br>same<br>? Like<br>nat w<br>eal so<br>ay no | s bal<br>enth<br>ct car<br>spee<br>e in th<br>ould p<br>ometh<br>ot wo                     | l to d<br>is so<br>nnot l<br><i>ed of 1</i><br><i>he</i><br><i>berming. E</i><br><i>rk if b</i> | irect<br>muc<br>be<br><i>a fla</i><br><i>m/s</i><br>t<br>But to<br>oth | tly<br>ch |       |   |
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| its not v<br>see effe<br>smaller<br>observe<br><i>Input C</i><br><i>table, cc</i><br><i>running</i><br><i>Thompse</i><br><i>electron</i><br><i>set it up</i> | visible. So<br>ect of this<br>than the<br>ed experim<br><b>DR Comm</b><br>build a high<br>along side<br>on experin<br>flow thro<br>moving v | o if w<br>s wav<br>ball<br>ment<br>hly ch<br>e, rev<br>emen<br>ugh i<br>with a | ve we<br>veleng<br>of ma<br>ally.<br><i>OR Ja</i><br><i>pargeo</i><br><i>real al</i><br><i>t. The</i><br><i>t. May</i><br><i>spee</i> | ere to<br>gth w<br>ass 1<br>Ans.<br>oke:<br>d elec<br>ny pa<br>e ball<br>vbe a<br>ed of 1 | con<br>ould<br>ogra<br>If this<br>tron<br>ttern<br>be m<br>patte<br>1m/s | duct<br>be u<br>am/s<br><i>s ball</i><br>beam<br>on a<br>ade o<br>ern on<br>may l | an exp<br>inrealis<br>dimer<br><i>was ro</i><br><i>aimed</i><br><i>photog</i><br><i>of a ma</i><br><i>o the pl</i><br><i>be a hu</i> | beriem<br>stic. T<br>isions.<br><i>Illing at</i><br><i>at it a</i><br><i>raphic</i><br><i>terial t</i><br><i>late ma</i><br><i>irdle.</i> T | ent o<br>The w<br>The<br><i>a cor</i><br><i>t the s</i><br><i>plate</i><br><i>plate</i><br><i>type th</i><br><i>tyr reve</i> | n thi<br>avel<br>effec<br>nstan<br>same<br>? Like<br>nat w<br>eal so<br>ay no | s bal<br>enth<br>ct car<br>spee<br>e in th<br>ould p<br>ometh<br>ot wo                     | l to d<br>is so<br>nnot l<br><i>ed of 1</i><br><i>he</i><br><i>berming. E</i><br><i>rk if b</i> | irect<br>muc<br>be<br><i>a fla</i><br><i>m/s</i><br>t<br>But to<br>oth | tly<br>ch |       |   |
| its not v<br>see effe<br>smaller<br>observe<br><i>Input C</i><br><i>table, cc</i><br><i>running</i><br><i>Thompse</i><br><i>electron</i><br><i>set it up</i> | visible. So<br>ect of this<br>than the<br>ed experim<br><b>DR Comm</b><br>build a high<br>along side<br>on experin<br>flow thro<br>moving v | o if w<br>s wav<br>ball<br>ment<br>hly ch<br>e, rev<br>emen<br>ugh i<br>with a | ve we<br>veleng<br>of ma<br>ally.<br><i>OR Ja</i><br><i>pargeo</i><br><i>real al</i><br><i>t. The</i><br><i>t. May</i><br><i>spee</i> | ere to<br>gth w<br>ass 1<br>Ans.<br>oke:<br>d elec<br>ny pa<br>e ball<br>vbe a<br>ed of 1 | con<br>ould<br>ogra<br>If this<br>tron<br>ttern<br>be m<br>patte<br>1m/s | duct<br>be u<br>am/s<br><i>s ball</i><br>beam<br>on a<br>ade o<br>ern on<br>may l | an exp<br>inrealis<br>dimer<br><i>was ro</i><br><i>aimed</i><br><i>photog</i><br><i>of a ma</i><br><i>o the pl</i><br><i>be a hu</i> | beriem<br>stic. T<br>isions.<br><i>Illing at</i><br><i>at it a</i><br><i>raphic</i><br><i>terial t</i><br><i>late ma</i><br><i>irdle.</i> T | ent o<br>The w<br>The<br><i>a cor</i><br><i>t the s</i><br><i>plate</i><br><i>plate</i><br><i>type th</i><br><i>tyr reve</i> | n thi<br>avel<br>effec<br>nstan<br>same<br>? Like<br>nat w<br>eal so<br>ay no | s bal<br>enth<br>ct car<br>spee<br>e in th<br>ould p<br>ometh<br>ot wo                     | l to d<br>is so<br>nnot l<br><i>ed of 1</i><br><i>he</i><br><i>berming. E</i><br><i>rk if b</i> | irect<br>muc<br>be<br><i>a fla</i><br><i>m/s</i><br>t<br>But to<br>oth | tly<br>ch |       |   |
| its not v<br>see effe<br>smaller<br>observe<br><i>Input C</i><br><i>table, cc</i><br><i>running</i><br><i>Thompse</i><br><i>electron</i><br><i>set it up</i> | visible. So<br>ect of this<br>than the<br>ed experim<br><b>DR Comm</b><br>build a high<br>along side<br>on experin<br>flow thro<br>moving v | o if w<br>s wav<br>ball<br>ment<br>hly ch<br>e, rev<br>emen<br>ugh i<br>with a | ve we<br>veleng<br>of ma<br>ally.<br><i>OR Ja</i><br><i>pargeo</i><br><i>real al</i><br><i>t. The</i><br><i>t. May</i><br><i>spee</i> | ere to<br>gth w<br>ass 1<br>Ans.<br>oke:<br>d elec<br>ny pa<br>e ball<br>vbe a<br>ed of 1 | con<br>ould<br>ogra<br>If this<br>tron<br>ttern<br>be m<br>patte<br>1m/s | duct<br>be u<br>am/s<br><i>s ball</i><br>beam<br>on a<br>ade o<br>ern on<br>may l | an exp<br>inrealis<br>dimer<br><i>was ro</i><br><i>aimed</i><br><i>photog</i><br><i>of a ma</i><br><i>o the pl</i><br><i>be a hu</i> | beriem<br>stic. T<br>isions.<br><i>Illing at</i><br><i>at it a</i><br><i>raphic</i><br><i>terial t</i><br><i>late ma</i><br><i>irdle.</i> T | ent o<br>The w<br>The<br><i>a cor</i><br><i>t the s</i><br><i>plate</i><br><i>plate</i><br><i>type th</i><br><i>tyr reve</i> | n thi<br>avel<br>effec<br>nstan<br>same<br>? Like<br>nat w<br>eal so<br>ay no | s bal<br>enth<br>ct car<br>spee<br>e in th<br>ould p<br>ometh<br>ot wo                     | l to d<br>is so<br>nnot l<br><i>ed of 1</i><br><i>he</i><br><i>berming. E</i><br><i>rk if b</i> | irect<br>muc<br>be<br><i>a fla</i><br><i>m/s</i><br>t<br>But to<br>oth | tly<br>ch |       |   |
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| its not v<br>see effe<br>smaller<br>observe<br><i>Input C</i><br><i>table, cc</i><br><i>running</i><br><i>Thompse</i><br><i>electron</i><br><i>set it up</i> | visible. So<br>ect of this<br>than the<br>ed experim<br><b>DR Comm</b><br>build a high<br>along side<br>on experin<br>flow thro<br>moving v | o if w<br>s wav<br>ball<br>ment<br>hly ch<br>e, rev<br>emen<br>ugh i<br>with a | ve we<br>veleng<br>of ma<br>ally.<br><i>OR Ja</i><br><i>pargeo</i><br><i>real al</i><br><i>t. The</i><br><i>t. May</i><br><i>spee</i> | ere to<br>gth w<br>ass 1<br>Ans.<br>oke:<br>d elec<br>ny pa<br>e ball<br>vbe a<br>ed of 1 | con<br>ould<br>ogra<br>If this<br>tron<br>ttern<br>be m<br>patte<br>1m/s | duct<br>be u<br>am/s<br><i>s ball</i><br>beam<br>on a<br>ade o<br>ern on<br>may l | an exp<br>inrealis<br>dimer<br><i>was ro</i><br><i>aimed</i><br><i>photog</i><br><i>of a ma</i><br><i>o the pl</i><br><i>be a hu</i> | beriem<br>stic. T<br>isions.<br><i>Illing at</i><br><i>at it a</i><br><i>raphic</i><br><i>terial t</i><br><i>late ma</i><br><i>irdle.</i> T | ent o<br>The w<br>The<br><i>a cor</i><br><i>t the s</i><br><i>plate</i><br><i>plate</i><br><i>type th</i><br><i>tyr reve</i> | n thi<br>avel<br>effec<br>nstan<br>same<br>? Like<br>nat w<br>eal so<br>ay no | s bal<br>enth<br>ct car<br>spee<br>e in th<br>ould p<br>ometh<br>ot wo                     | l to d<br>is so<br>nnot l<br><i>ed of 1</i><br><i>he</i><br><i>berming. E</i><br><i>rk if b</i> | irect<br>muc<br>be<br><i>a fla</i><br><i>m/s</i><br>t<br>But to<br>oth | tly<br>ch |       |   |
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| its not v<br>see effe<br>smaller<br>observe<br><i>Input C</i><br><i>table, cc</i><br><i>running</i><br><i>Thompse</i><br><i>electron</i><br><i>set it up</i> | visible. So<br>ect of this<br>than the<br>ed experim<br><b>DR Comm</b><br>build a high<br>along side<br>on experin<br>flow thro<br>moving v | o if w<br>s wav<br>ball<br>ment<br>hly ch<br>e, rev<br>emen<br>ugh i<br>with a | ve we<br>veleng<br>of ma<br>ally.<br><i>OR Ja</i><br><i>pargeo</i><br><i>real al</i><br><i>t. The</i><br><i>t. May</i><br><i>spee</i> | ere to<br>gth w<br>ass 1<br>Ans.<br>oke:<br>d elec<br>ny pa<br>e ball<br>vbe a<br>ed of 1 | con<br>ould<br>ogra<br>If this<br>tron<br>ttern<br>be m<br>patte<br>1m/s | duct<br>be u<br>am/s<br><i>s ball</i><br>beam<br>on a<br>ade o<br>ern on<br>may l | an exp<br>inrealis<br>dimer<br><i>was ro</i><br><i>aimed</i><br><i>photog</i><br><i>of a ma</i><br><i>o the pl</i><br><i>be a hu</i> | beriem<br>stic. T<br>isions.<br><i>Illing at</i><br><i>at it a</i><br><i>raphic</i><br><i>terial t</i><br><i>late ma</i><br><i>irdle.</i> T | ent o<br>The w<br>The<br><i>a cor</i><br><i>t the s</i><br><i>plate</i><br><i>plate</i><br><i>type th</i><br><i>tyr reve</i> | n thi<br>avel<br>effec<br>nstan<br>same<br>? Like<br>nat w<br>eal so<br>ay no | s bal<br>enth<br>ct car<br>spee<br>e in th<br>ould p<br>ometh<br>ot wo                     | l to d<br>is so<br>nnot l<br><i>ed of 1</i><br><i>he</i><br><i>berming. E</i><br><i>rk if b</i> | irect<br>muc<br>be<br><i>a fla</i><br><i>m/s</i><br>t<br>But to<br>oth | tly<br>ch |       |   |

|                                       | e Broglie wavelength of thermal neutrons at 27 deg C.<br>n constant k = $1.38 \times 10^{-23}$ J/K. |  |
|---------------------------------------|---|--|
| Solution:                             |   |  |
| $m_{neutron} := 1.67 \cdot T := 27 C$ | • 10 <sup>-27</sup> kg  |  |
| $T_{K} \coloneqq T + 273 = 3$         | 300 К.  |  |
| k <sub>Boltzman</sub> ≔1.38           | • 10 <sup>-23</sup> J/K.  |  |
| KE_thermal_n                          | $eutron := k_{Boltzman} \cdot T_{K} = 4.14 \cdot 10^{-21}$  |  |
| λ:=                                   | (h) = $1.775 \cdot 10^{-10}$  |  |
| $\sqrt{2 \cdot m_{neutro}}$           | $\frac{(h)}{(h)} = 1.775 \cdot 10^{-10}$  |  |
| $\lambda \coloneqq 1.78 \ Angst$      | rom Ans.  |  |
| Problem 4.8                           |   |  |
| •                                     | eutron have the same kinetic energy.<br>vo has longer de Broglie wavelength?                        |  |
| Solution:                             |   |  |
|                                       | e, 2X, the mass of proton.<br>e of interest in solving this example problem.                        |  |
| m-p: mass of pr                       |   |  |
| m-d: mass of de $m-d = 2 x (m-p)$     |   |  |
| lamda-p : h/ Sq                       | ırt (2 m-р K)   |  |
| lamda_d : h/ Sc                       | qrt (2 (2 X m-d) K)   |  |
| lamda_p / lamd                        | la_d = h Sqrt (2 (2 X m-d) K)<br>Sqrt (2 m-p K) h   |  |
|                                       | la_d = Sqrt(2) Ans.   |  |

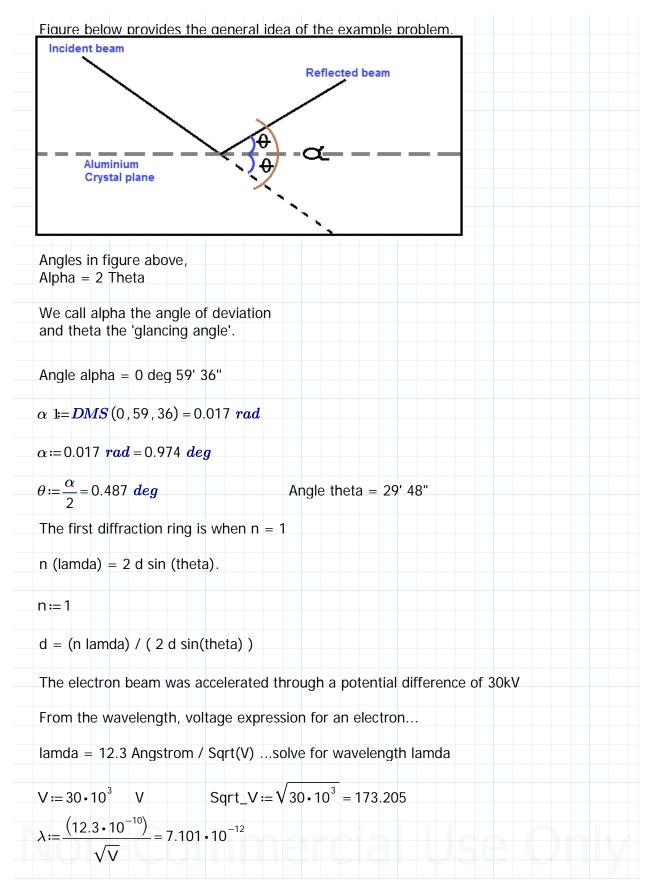
|        | that the de Broglie wavelength of an electron is equal to its compton wavelength its speed is <u>c/sqrt(2)</u> , c being the speed of light. |
|--------|--|
| Soluti | on:  |
| Comp   | ton wavelength = $h/((m_0) c)$ .   |
| de Br  | Is known as the Compton wavelength of the electron.<br>oglie wavelength = $h / p = h / mv = h / Sqrt (2 m K)$ .                              |
| The e  | expressions which are smilar among the two are   |
| Comp   | oton> <u>h/ ((m_0) c)</u> and <u>h /mv</u> < Broglie   |
|        | out going into the details on the background theory, and mathematical  |
| detail | s, the equation<br>$E = mc^2 = (m_0 (c^2)) / Sqrt(1 - (v^2/c^2))$  |
| m_0    | is the rest mass of a particle and v its velocity, and   |
|        | $m_0 / Sqrt(1 - (v^2/c^2))$  |
|        | ained through Einstein's work in change of mass of a body with velocity. This  |
|        | e studied in any advanced Physics or most Modern Physics textbook. <i>One</i><br><i>pok referenced here is SN Ghosal Modern Physics.</i>     |
| lamda  | a = h / (mv)de Broglie   |
|        | $m = m_0 / Sqrt(1 - (v^2/c^2))$  |
|        | = $(h/m_0 v) [Sqrt(1 - (v^2/c^2))]$  |
| subst  | itute v for c multiply by (c/v)keeping things the same   |
| lamda  | $a = (h/m_0 c) [(1 - (v^2/c^2))^{(1/2)}] (c/v) = h/(mv) \dots Correct.$  |
| subst  | itute v (velocity) = c/Sqrt(2).  |
| lamda  | a = (h/m_0 c) [( 1 - (c/Sqrt(2)^2/c^2) )^(1/2)] (c/(c/Sqrt(2))) = h/(mv)   |
| lamda  | a = (h/m_0 c) [( 1 - ((c^2/2)/c^2) )^(1/2)] (c/(c/Sqrt(2))) = h/(mv)   |
| lamda  | a = (h/m_0 c) [( 1 - (1/2) )^(1/2)] (Sqrt(2)) = h/(mv)   |
|        | a = (h/m_0 c) [( 1/2)^(1/2)] (Sqrt(2)) = h/(mv)  |
|        | $a = (h/m_0 c) [Sqrt(1/2)] (Sqrt(2)) = h/(mv)$   |
|        | $a = (h/m_0 c) (1) = h/(mv)$   |
| anua   | a = (h/m_0 c) is the Compton wavelength of the electron. Ans.  |

Quantum Mechanics A Textbook for Undergraduates. Mahesh C. Jain. PHI. Department of Physics, Hindu College University of Delhi. To Support Studies In: Modern Physics by S.N. Ghosal. Purpose: Quantum Mechanics for 'Power Plant Engineering' Studies. Exercise by: K S Bogha. Part III Wave Nature of Matter - de Broglie. Problem 4.10 In a Davidson-Germer diffraction experiment electrons of kinetic energy 100 eV are scattered from a crystal. The first maximum in intensity occurs at theta = 10.0 deg. (a). What is the spacing between the crystal planes? (b). How many peaks will there be in the interference pattern? Solution: See Modern Physics textbooks for Davisson Germer experiement and Bragg planes for crystals. The formula  $\phi_{/_2}$ provided here their deriviation  $\phi_{/_2}$ Reflected beam ð are provided there. Figure to the right shows the Ð angles concerned in the Diffraction of electron equation. waves by a crystal Bragg planes Basic derivation for the equation to apply provided below. ?Ø/2 Øђ a). n (lamda) = 2 d sin (theta)d is the spacing between Bragg planes, and n is an integer. Angle theta and phi are shown in the figure above. (theta) + (phi/2 + phi/2) + (theta) = 180 degsor (theta) + (phi) + (theta)= 180 degs (theta) = (180 degs - phi) / 2(theta) = 90 degs - (phi / 2)Applying geometry, and trignometry sin(phi/2) = d/D .....where D is the interatomic spacing.  $d = D \sin(phi/2)$ 

| n(lamda) = 2 D sin (ph                                   |  |
|--|--|
|  | sin (theta)  |
| n(lamda) = 2 D sin (ph                                   | 11/2) cos (pni/2)  |
| tria ide   | ntity cos (phi/2) = sin (90 - (phi/2))                             |
| trig identity sin 2(phi)                                 |  |
|  | 2 sin(phi/2)cos(phi/2)   |
| subisituting into 2 D sir                                | n (phi/2) cos (phi/2) and the 2 cancels.                           |
| n(lamda) = D sin (phi)                                   | this is the equation to be applied.                                |
| V:=100 V   |  |
| $123.10^{-10}$   |  |
| $\lambda \coloneqq \frac{12.3 \cdot 10^{-10}}{\sqrt{V}}$ | expression for electron.   |
| V  |  |
| $\lambda = 1.23 \cdot 10^{-10}$                          |  |
| $\lambda = 1.23 \cdot 10$                                |  |
| <i>θ</i> ≔10 <i>deg</i>                                  |  |
|  |  |
| $d := \frac{\lambda}{2 \cdot \sin(\theta)} = 3.54$       | ·2 • 10 <sup>-10</sup>   |
| $2 \cdot \sin(\theta)$                                   |  |
|  |  |
| d≔3.542 Angstro  | m Ans.   |
| b).  |  |
| n (lamda) = 2 d sin (th)                                 | leta)  |
|  |  |
|  | , and the variable is sin(theta) for the maximum value of          |
| n sin(theta) has to be r                                 |  |
| $\sin(\tan \theta a) = 1$ is a positive                  | sible maximum when angle theta = 90 degs                           |
| now the expression be                                    | comes:   |
|  |  |
| n (lamda) = 2 d sin (90 $r$                              | ))   |
| n (lamda) = 2 d (1)<br>n = $2d/lamda$                    |  |
|  | annot be greater then the RHS so                                   |
| n < OR = 2d/lamda  |  |
| $n := \frac{2 \cdot d}{10} = 5,759.10^{10}$              | Therefore the maximum value n can take is 5, it cannot be larg     |
| $n := \frac{2 \cdot d}{\lambda} = 5.759 \cdot 10^{10}$   | then the next integer 6.   |
|  | Comment: PV panel has a thickness and so do the PV cells, so could |

| of 30 kV, pass   | im of electrons, accelerated thro<br>ses through a thin aluminium foil<br>photographic plate on the opposi | and produces a diffraction   |
|--|--|--|
|  | raction ring is obtained at an an<br>, calculate the grating space in t                                    |  |
| Solution:  |  |  |
| You are welco<br>said in the ske                       |  | eral idea.<br>t this experiment portray(s). What I have<br>u. For now to solve the example problem   |
|  | 1  | Photographic plate   |
| <u> </u>   | hompson's experiment.  |  |
|  | Aluminium Crystals   | Series of concentric<br>rings appeared<br>proving de Broglie<br>hypothesis.<br>Do the rings identify to<br>electron rings?<br>Yes. diffraction pattern of  |
| Incident beam<br>of highly<br>accelerated<br>electrons | Aluminium Crystals   | rings appeared<br>proving de Broglie<br>hypothesis.<br>Do the rings identify to<br>electron rings?<br>Yes. diffraction pattern of<br>electrons.<br>de Broglie awarded<br>Noble Prize for Physics<br>in 1929. In 1936 Davisson<br>and Thompson shared the |
| Incident beam<br>of highly<br>accelerated              | Aluminium Crystals   | rings appeared<br>proving de Broglie<br>hypothesis.<br>Do the rings identify to<br>electron rings?<br>Yes. diffraction pattern of<br>electrons.<br>de Broglie awarded<br>Noble Prize for Physics<br>in 1929. In 1936 Davisson                            |
| Incident beam<br>of highly<br>accelerated              | Aluminium Crystals   | rings appeared<br>proving de Broglie<br>hypothesis.<br>Do the rings identify to<br>electron rings?<br>Yes. diffraction pattern of<br>electrons.<br>de Broglie awarded<br>Noble Prize for Physics<br>in 1929. In 1936 Davisson<br>and Thompson shared the |

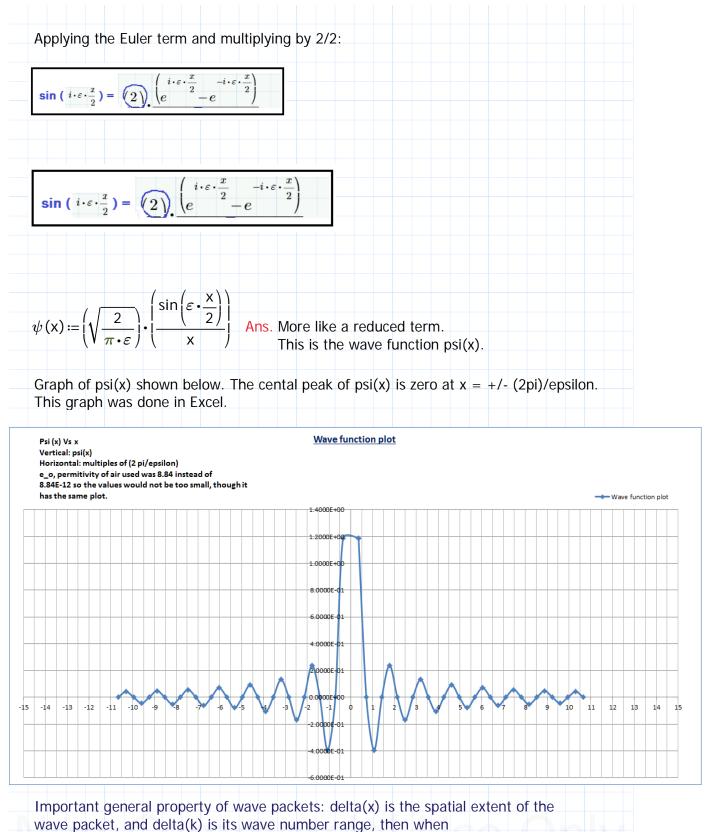
**FYI:** This experiment was done by the son, the father also a physicist was JJ Thompson, he (father) received a Nobel Prize for showed electron has a particle nature in 1906. The son GP Thompson received a Nobel Prize in 1937 shared with Davisson, showed the electron has a wave nature. You verify.



| $Sin_{theta} = sin(\theta) = 0.0085$   |        |
|--|--------|
| Sin_theta:=0.0087  |        |
| $d \coloneqq \frac{n \cdot \lambda}{2 \cdot \sin(\theta)}$   |        |
| $d = 4.177 \cdot 10^{-10}$   |        |
| d := 4.177 $Angstrom$ Ans.   |        |
| Part III basically a few examples using formulas that are short not l involved. Introductory level.  | nighly |
| Any errors and omissions apologies in advance.   |        |
| Thus far the 3 part solar engineering tutorials and partially study material<br>may be sufficient to take you to the start of Wave Packets and The Uncern<br>Principle topics. This topics require a textbook to master the theories, mat<br>deriviations, and related subject matter. | tainty |
| Requires you to have a Modern Physics textbook for the study material.   |        |
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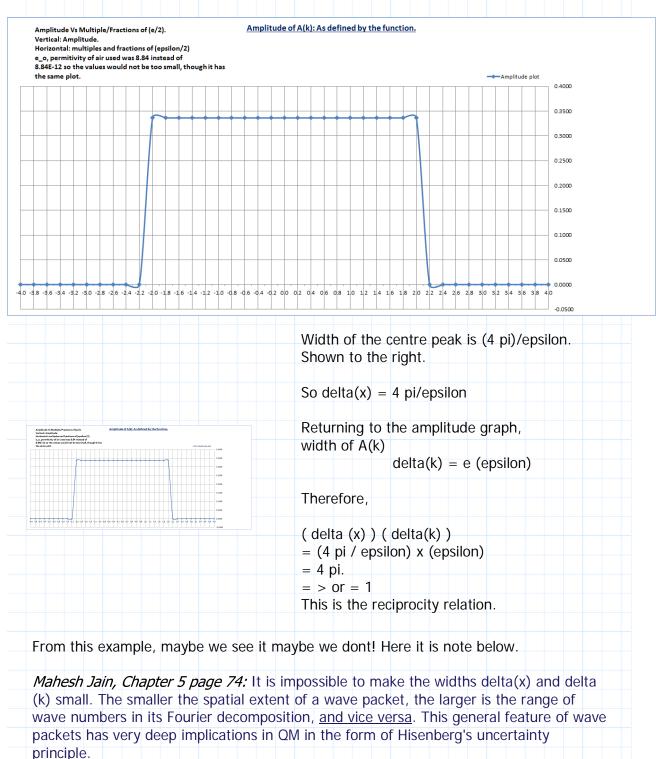
|   | tbook: Modern (Atomic) Physics Vol I.  |
|---|--|
|   | hor:S.N. Ghosal  |
| Put   | lisher: S. Chand.  |
| Cha   | pter 9: Wave particle duality; Hisenberg's uncertainty principle.  |
| 9.1   | Particle and waves   |
| 9.2   | Phase and group velocities   |
| 9.3   |  |
| 9.4   | Relation between phase and group velocity of de Broglie waves  |
| 9.5   |  |
| 9.6   | G.P. Thompson's experiment   |
| 9.7   | Effect of refraction of the electron beam  |
| 9.8   | De Broglie wavelength of high energy electrons   |
| 9.9   | Electron microscope  |
|   | D Need for a new mechanics for the sub-atomic particles  |
| 9.1   | 1 Particles and wave packets   |
|   | 2 Nature of matter waves   |
|   | 3 Uncertainty relations  |
|   | 4 Gamma-ray microscope experiment  |
|   | 5 Applicability of classical and quantum concepts  |
| 9.1   | 6 Principle of superposition   |
|   | nesh C Jain.   |
| 5.1   | Representation of a particle by wave packet  |
|   | Heisenberg's Uncertainty Principle   |
|   | Illustrations of the Uncertainty Principle   |
| 5.3   |  |
| 5.3   | Application/Consequences of the Uncertainty Principle  |
| 5.3<br>5.4  | Application/Consequences of the Uncertainty Principle  |
| 5.3<br>5.4<br><b>Exa</b>  | Application/Consequences of the Uncertainty Principle<br>mple solved problems here are mostly on the red text above (Mahesh).  |
| 5.3<br>5.4<br>Exa<br>Thi  | Application/Consequences of the Uncertainty Principle  |
| 5.3<br>5.4<br>Exa<br>Thi<br>sut   | Application/Consequences of the Uncertainty Principle<br>mple solved problems here are mostly on the red text above (Mahesh).<br>s for me is the hardest part in the early material in an UG quantum mechanics course. The<br>ject matter on wave mechanics. <u>Problems here more on how to apply some relations</u> .<br>are is no need for notes here because the subject matter is best studied in modern physics  |
| 5.3<br>5.4<br>Exa<br>Thi<br>sut   | Application/Consequences of the Uncertainty Principle<br>mple solved problems here are mostly on the red text above (Mahesh).<br>s for me is the hardest part in the early material in an UG quantum mechanics course. The<br>ject matter on wave mechanics. <u>Problems here more on how to apply some relations.</u>   |
| 5.3<br>5.4<br>Exa<br>Thi<br>sub<br>The<br>or (                            | Application/Consequences of the Uncertainty Principle<br>mple solved problems here are mostly on the red text above (Mahesh).<br>s for me is the hardest part in the early material in an UG quantum mechanics course. The<br>ject matter on wave mechanics. Problems here more on how to apply some relations.<br>ere is no need for notes here because the subject matter is best studied in modern physics<br>quantum mechanics textbooks. Certainly in quantum mechanics textbooks.  |
| 5.3<br>5.4<br>Exa<br>Thi<br>suk<br>The<br>or (                            | Application/Consequences of the Uncertainty Principle<br>mple solved problems here are mostly on the red text above (Mahesh).<br>s for me is the hardest part in the early material in an UG quantum mechanics course. The<br>ject matter on wave mechanics. Problems here more on how to apply some relations.<br>ere is no need for notes here because the subject matter is best studied in modern physics<br>quantum mechanics textbooks. Certainly in quantum mechanics textbooks.<br>e relationship between phase and group velocities need to be understood in theory,<br>thematical explanation and derivation. This relationship can be shown using Fourier Series  |
| 5.3<br>5.4<br>Exa<br>Thi<br>suk<br>The<br>or (                            | Application/Consequences of the Uncertainty Principle<br>mple solved problems here are mostly on the red text above (Mahesh).<br>s for me is the hardest part in the early material in an UG quantum mechanics course. The<br>ject matter on wave mechanics. Problems here more on how to apply some relations.<br>ere is no need for notes here because the subject matter is best studied in modern physics<br>quantum mechanics textbooks. Certainly in quantum mechanics textbooks.  |
| 5.3<br>5.4<br>Exa<br>Thi<br>suk<br>The<br>or o<br>The<br>ma<br>Tra        | Application/Consequences of the Uncertainty Principle<br>mple solved problems here are mostly on the red text above (Mahesh).<br>s for me is the hardest part in the early material in an UG quantum mechanics course. The<br>ject matter on wave mechanics. Problems here more on how to apply some relations.<br>ere is no need for notes here because the subject matter is best studied in modern physics<br>quantum mechanics textbooks. Certainly in quantum mechanics textbooks.<br>e relationship between phase and group velocities need to be understood in theory,<br>thematical explanation and derivation. This relationship can be shown using Fourier Series<br>nsforms, OR Algebraic-Differentiation deriviations. So no doubt its a little involved.    |
| 5.3<br>5.4<br>Exa<br>Thi<br>sut<br>The<br>or o<br>The<br>ma<br>Tra<br>Hav | Application/Consequences of the Uncertainty Principle<br>mple solved problems here are mostly on the red text above (Mahesh).<br>s for me is the hardest part in the early material in an UG quantum mechanics course. The<br>ject matter on wave mechanics. Problems here more on how to apply some relations.<br>there is no need for notes here because the subject matter is best studied in modern physics<br>quantum mechanics textbooks. Certainly in quantum mechanics textbooks.<br>e relationship between phase and group velocities need to be understood in theory,<br>thematical explanation and derivation. This relationship can be shown using Fourier Series<br>insforms, OR Algebraic-Differentiation deriviations. So no doubt its a little involved. |
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| 5.3<br>5.4<br>Exa<br>Thi<br>sub<br>The<br>ma<br>Tra<br>Hay<br>and         | Application/Consequences of the Uncertainty Principle<br>mple solved problems here are mostly on the red text above (Mahesh).<br>s for me is the hardest part in the early material in an UG quantum mechanics course. The<br>ject matter on wave mechanics. Problems here more on how to apply some relations.<br>there is no need for notes here because the subject matter is best studied in modern physics<br>quantum mechanics textbooks. Certainly in quantum mechanics textbooks.<br>e relationship between phase and group velocities need to be understood in theory,<br>thematical explanation and derivation. This relationship can be shown using Fourier Series<br>insforms, OR Algebraic-Differentiation deriviations. So no doubt its a little involved. |

| h ≔ 6.63 • 10   | JS  |   |
|---|---|---|
| $c := 3 \cdot 10^8$<br>eV := 1.6 • 10   | m/s<br>$J^{-19}$ J<br>$J^{19}$ J<br>$J^{-27}$ $J^{-27}$   |   |
| eV≔1.6•10   | ) <sup>-19</sup> J  |   |
| $e = 1.6 \cdot 10^{-1}$   | 19 J  |   |
| m <sub>neutron</sub> ≔ 1  | .675•10 <sup>-27</sup> kg   |   |
| $m_{neutron} := 1$<br>$m_{electron} := \langle $  | 9.1•10 <sup>-31</sup> ) kg  |   |
| Problem 5.1   | _   |   |
| A wave pac  | ket has the amplitu   | de function   |
| 1/so  | rt(epsilon), - epsilo   | on/2 <= k <= epsilon/2  |
| A(k) =  |   |   |
| 0,  | k  > 6  | epsilon/2   |
| Find the wa   | ve function psi(x) a  | and hence verify the reciprocity relation   |
|   |   | nt know Fourier dont worry just go thru the steps and the no  |
| Look up on  | the internet just ge  | et a general idea of its purpose.   |
| Solution:   |   |   |
| SUIUTION  |   |   |
| $i := \sqrt{-1}$  | $\epsilon_{0} = 8.84 \cdot 10^{-12}$  | F/m farad per meter here its permitivity of air,  |
| 1. <u> </u>   | 20-0.04•10  | just picked it so it does not show a red flag in the  |
| <i>ε</i> ≔8.84  |   | expressionerror in software data entry.   |
| 2 - 0.04  |   |   |
| nsi(x) term   | is obtained by takin  | ng the Fourier transform of A(k):   |
|   |   |   |
|   |   |   |
|   | ∞<br>1 \ <b>€</b>   | Short Talk on the side:   |
| $\psi(x) \coloneqq \left( $   | $\left \frac{1}{k}\right  \cdot \int_{-\infty}^{\infty} A(k) \cdot e^{i \cdot k}$   | Short Talk on the side:   |
| $\psi(\mathbf{x}) \coloneqq \left(\frac{1}{\sqrt{2}}\right)$  | $\frac{1}{1-\pi} \cdot \int_{-\infty}^{\infty} A(k) \cdot e^{i \cdot k}$  | The A(k) is function but more like the  |
| $\psi(\mathbf{x}) \coloneqq \left(\frac{1}{\sqrt{2}}\right)$  | $\frac{1}{1 \cdot \pi} \cdot \int_{-\infty}^{\infty} \mathbf{A}(k) \cdot e^{i \cdot k}$   | The A(k) is function but more like the conditions on how works, where it  |
|   | ζ   | The A(k) is function but more like the conditions on how works, where it values rests for non-zero and zero.  |
|   | ζ   | The A(k) is function but more like the conditions on how works, where it values rests for non-zero and zero.  |
|   | ζ   | The A(k) is function but more like the conditions on how works, where it values rests for non-zero and zero.  |
|   | ζ   | The A(k) is function but more like the conditions on how works, where it values rests for non-zero and zero.  |
|   | ζ   | The A(k) is function but more like the conditions on how works, where it values rests for non-zero and zero.  |
| $\psi(\mathbf{x}) \coloneqq \left(\frac{1}{\sqrt{2}}\right)$  | $\frac{1}{1 \cdot \pi} \cdot \int_{\frac{-\varepsilon}{2}}^{\frac{\varepsilon}{2}} \left(\frac{1}{\sqrt{\varepsilon}}\right) \cdot e^{i \cdot \frac{\varepsilon}{2}}$   | *** dx       The A(k) is function but more like the conditions on how works, where it values rests for non-zero and zero.         *** dx       So by applying some vague or super accurate method we try to get a function on A(k) that takes the form of a wave. Thats all!  |
| $\psi(\mathbf{x}) \coloneqq \left(\frac{1}{\sqrt{2}}\right)$  | $\frac{1}{1 \cdot \pi} \cdot \int_{\frac{-\varepsilon}{2}}^{\frac{\varepsilon}{2}} \left(\frac{1}{\sqrt{\varepsilon}}\right) \cdot e^{i \cdot \frac{\varepsilon}{2}}$   | *** dx       The A(k) is function but more like the conditions on how works, where it values rests for non-zero and zero.         *** dx       So by applying some vague or super accurate method we try to get a function on A(k) that takes the form of a wave. Thats all!  |
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| $\psi(\mathbf{x}) \coloneqq \left(\frac{1}{\sqrt{2}}\right)$ $\psi(\mathbf{x}) = \left(\frac{1}{\sqrt{2}}\right)$ | $\frac{1}{2\cdot\pi}\right) \cdot \int_{-\frac{\varepsilon}{2}}^{\frac{\varepsilon}{2}} \left(\frac{1}{\sqrt{\varepsilon}}\right) \cdot e^{i\cdot}$ $\frac{1}{2\cdot\pi\cdot\varepsilon}\right) \cdot \left(\frac{e^{i\cdot k\cdot x}}{i\cdot x}\right)^{\frac{\varepsilon}{2}}$ $-\frac{\varepsilon}{2}$ | <ul> <li><sup>k · x</sup> dx</li> <li>The A(k) is function but more like the conditions on how works, where it values rests for non-zero and zero.</li> <li><sup>k · x</sup> dx</li> <li>So by applying some vague or super accurate method we try to get a function on A(k) that takes the form of a wave. Thats all!</li> <li>Which came first the math or the signal theory? Math. Fourier math came first and it was worthy to apply in signals which signals you know are</li> </ul> |
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| $\psi(\mathbf{x}) \coloneqq \left(\frac{1}{\sqrt{2}}\right)$ $\psi(\mathbf{x}) = \left(\frac{1}{\sqrt{2}}\right)$ | $\frac{1}{2\cdot\pi}\right) \cdot \int_{-\frac{\varepsilon}{2}}^{\frac{\varepsilon}{2}} \left(\frac{1}{\sqrt{\varepsilon}}\right) \cdot e^{i\cdot}$ $\frac{1}{2\cdot\pi\cdot\varepsilon}\right) \cdot \left(\frac{e^{i\cdot k\cdot x}}{i\cdot x}\right)^{\frac{\varepsilon}{2}}$ $-\frac{\varepsilon}{2}$ | <ul> <li><sup>k · x</sup> dx</li> <li>The A(k) is function but more like the conditions on how works, where it values rests for non-zero and zero.</li> <li><sup>k · x</sup> dx</li> <li>So by applying some vague or super accurate method we try to get a function on A(k) that takes the form of a wave. Thats all!</li> <li>Which came first the math or the signal theory? Math. Fourier math came first and it was worthy to apply in signals which signals you know are</li> </ul> |
| $\psi(\mathbf{x}) \coloneqq \left(\frac{1}{\sqrt{2}}\right)$ $\psi(\mathbf{x}) = \left(\frac{1}{\sqrt{2}}\right)$ | $\frac{1}{1 \cdot \pi} \cdot \int_{\frac{-\varepsilon}{2}}^{\frac{\varepsilon}{2}} \left(\frac{1}{\sqrt{\varepsilon}}\right) \cdot e^{i \cdot \frac{\varepsilon}{2}}$   | <ul> <li><sup>k · x</sup> dx</li> <li>The A(k) is function but more like the conditions on how works, where it values rests for non-zero and zero.</li> <li><sup>k · x</sup> dx</li> <li>So by applying some vague or super accurate method we try to get a function on A(k) that takes the form of a wave. Thats all!</li> <li>Which came first the math or the signal theory? Math. Fourier math came first and it was worthy to apply in signals which signals you know are</li> </ul> |



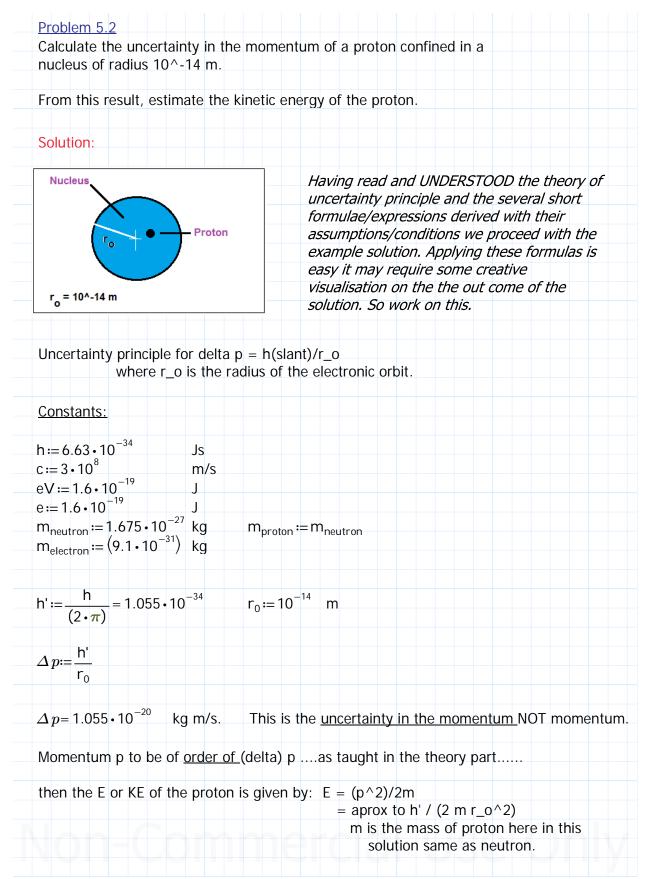
( delta (x) ) ( delta (k) ) >=1, reciprocity relation exist.

The graph of the amplitude A(K), is between (-e/2) and (e/2). In Excel the edges are not exactly squared off. Could be done in Mathcad, instead Excel used for plot and tabulation.



So smaller delta(x) larger delta(K), and larger delta(x) smaller delta(k). We show its product is greater than or equal 1 then it has reciprocity.

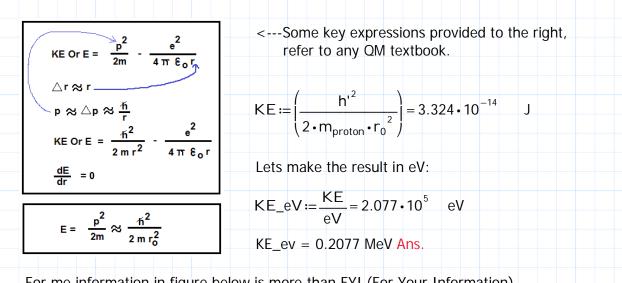
| tum Mechanics for 'Power Plant Engineering' Studies.<br>S Bogha. Part IV <u>Wave Packets and The Uncertainty Principle.</u>                                     |
|---|
|   |
| Reciprocity in mathematics is variable A multiplied by B results in 1.  |
| If $A = 3/4$ then B has to equal $4/3$ .  |
| Here delta(x) x delta(k) = 1 is a reciprocity   |
| plus we extend it to values greater than 1.   |
|   |
| 'In QM a particle, electron-neutron, is represented by a wave packet.   |
| The wave packet is made up of multiple waves, each wave has a phase velocity.   |
| Each packet has one group velocity.   |
| So can we determine where the particle represented by a wave packet is going  |
| to hit within a given area of choice? No. You may be lucky sometimes but in an  |
| empirical sense No.   |
| That No is part of a foundation on the Uncertainty Principle.' - My Wording.  |
| Mahesh puts it as 'the position of the particle is indeterminate within the width of  |
| the wave packet. Similarly the momentum of the particle is indeterminate within   |
| the region where the momentum wave function psi(p) is non zero.   |
|   |
| Momentum is mass x velocity, if the velocity is unable to precisely locate where  |
| the particle will land/strike on the choice area, no less can the momentum.   |
| We ask why is it so critical to know where the particle lands?  |
| Take the solar panels for example, an easy case, we want the photon to land on the  |
| electron hit it hard throw the electron out of orbit and start flowing. So in this perspective  |
| knowing the location and precision of strike or landing may be important. We dont make  |
| ONE solar panel (PV) the size of a football field! You agree? Goal!   |
|   |
| Returning to reciprocity, the x is distance or coordinate in the x-axis, k the space wave   |
| function, here $delta(x) \times delta(k) > or = 1$ .  |
| We know h' multiplied by k results in p   |
| h' = h/(2 pi)   |
| h'k = p   |
| k = p / h'  |
| substitute k into the reciprocity expression,   |
| delta(x) delta(p) > or = h.   |
|   |
| Ok. But does this give us any better precision than before? No.   |
|   |
| However it extends our lack of precision and turns into an understanding of   |
| uncertainty. Hisenberg's uncertainty principle states: It is not possible to specify both the position and momentum of a particle simultaneously with artibrary |
| precision; the product of uncertainties in the position and momentum is always  |
| greater than a quantity of order h' (ie h/(2 pi)) - Mahesh.   |



Quantum Mechanics A Textbook for Undergraduates. Mahesh C. Jain. PHI. Department of Physics, Hindu College University of Delhi. To Support Studies In: Modern Physics by S.N. Ghosal.

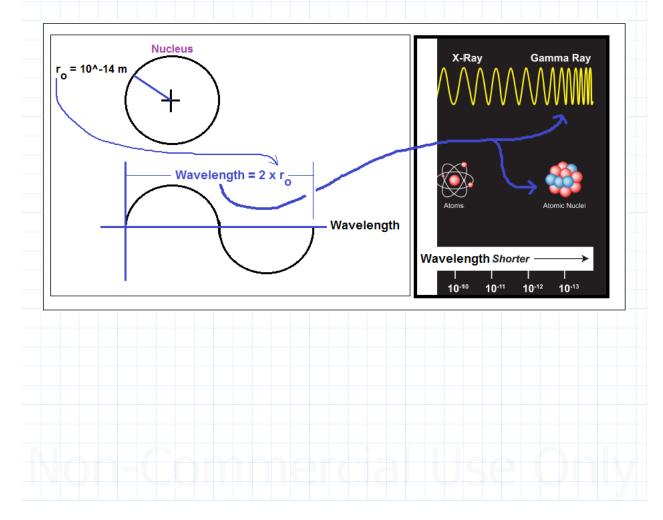
Purpose: <u>Quantum Mechanics for 'Power Plant Engineering' Studies.</u>

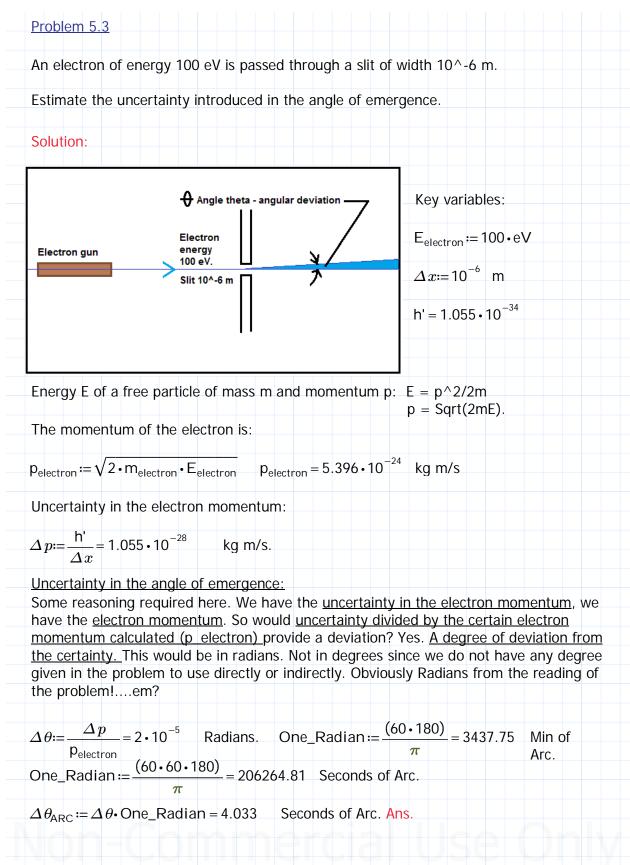
Exercise by: K S Bogha. Part IV Wave Packets and The Uncertainty Principle.

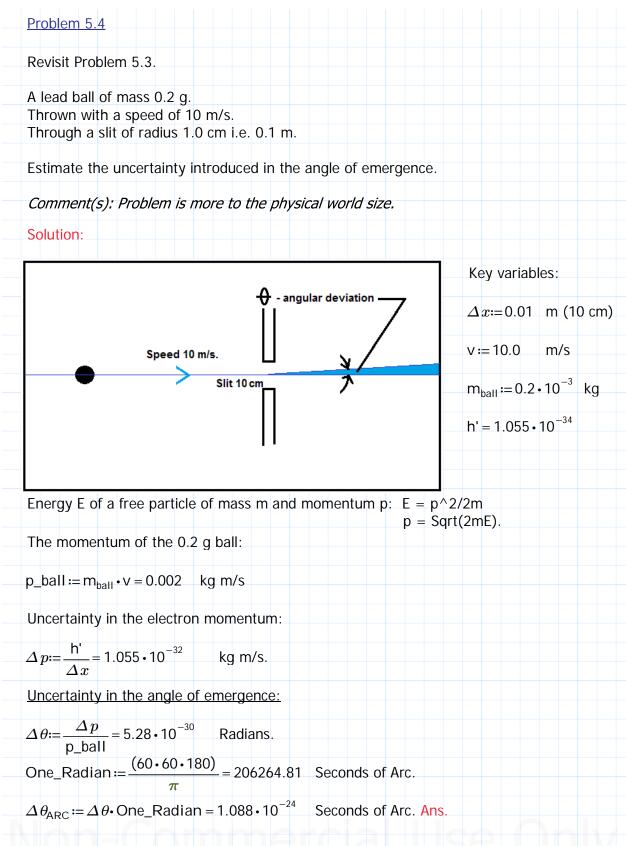


For me information in figure below is more than FYI (For Your Information). I am not a Physics major hence relating wavelength to the spectrum is not typical work for me, for most electrical engineers in communications it is!

So the wavelength is around the Gamma Ray range...which our problem is close too.







| Problem 5.5   |   |
|---|---|
| Speed of a bullet of mass 50 g is measured to be 30 Measured at an accuracy of 0.01%.   | )0 m/s.   |
| What accuracy can we locate the position of the bull  | let?  |
| Comment(s): A slight twist from the past 2 problems   | s, to find the location (delta)x.   |
| Solution:   |   |
|   | Key variables:  |
|   | v:=300.0 m/s  |
| Mass: 50 g  | v <sub>accuracy</sub> ≔0.01% Perce  |
|   | $m_{bullet} := 50 \cdot 10^{-3}$  |
| Accuracy of Measurement: 0.01%.   | $h' = 1.055 \cdot 10^{-34}$   |
|   |   |
| Uncertainty in momentum (delta)p = (delta) mv = m   | n (delta) v   |
| Uncertainty in momentum (delta)p = (delta) mv = m<br>$v_{adjusted} := v \cdot v_{accuracy} = 0.03$ m/s.   | n (delta) v   |
|   | n (delta) v   |
| $v_{adjusted} \coloneqq v \cdot v_{accuracy} = 0.03$ m/s.   |   |
| $v_{adjusted} := v \cdot v_{accuracy} = 0.03 \text{ m/s.}$ $\Delta p := m_{bullet} \cdot v_{adjusted} = 0.002 \text{ kg m/s}$ $(delta)p (delta)x = h' \dots h' = h/(2 \text{ pi}) \dots and (delta)$ $\Delta x := \frac{h'}{h} = 7.035 \cdot 10^{-32} \text{ m Ans. This is the accurace}$                            | Ita)x is what we seek.<br>cy to locate the position.<br>mall compared to the other real world |
| $v_{adjusted} := v \cdot v_{accuracy} = 0.03 \text{ m/s.}$ $\Delta p := m_{bullet} \cdot v_{adjusted} = 0.002 \text{ kg m/s}$ $(delta)p (delta)x = h' \dots h' = h/(2 \text{ pi}) \dots and (delta)$ $\Delta x := \frac{h'}{\Delta p} = 7.035 \cdot 10^{-32} \text{ m Ans. This is the accurace}$ It is remarkably so | Ita)x is what we seek.<br>cy to locate the position.<br>mall compared to the other real world |

| Problem 5.6                               |   |                                 |  |   |
|---|---|---------------------------------|--|---|
| The lifetime of                           | a nucleus in an excited   | l state is 10^-12s.             |  |   |
| Calculate the <u>pr</u><br>photon emitted | _   | the energy and free             | quency of a gamma ray                          |   |
| Comment(s): T                             | his can be an energy p  | oroblem in nuclear <sub>l</sub> | power plants.                                  |   |
| Solution:                                 |   |                                 |  |   |
| merely tries to                           | tches below may not b<br>give a sequence of eve<br>relative to particle (mo | ents and location of            | f the gamma ray in the                         |   |
| Molecules<br>Visible Light                | Atoms   |                                 | e from Ball Aerospace<br>nologies Corporation. |   |
| 10 <sup>-7</sup> 10 <sup>-8</sup> 1       | 10 <sup>-9</sup> 10 <sup>-10</sup> 10 <sup>-11</sup> 10 <sup>-12</sup>      | 10-13                           |  | 7 |
| Energy<br>level                           |   | Excited<br>state                | Emitted<br>gamma ray<br>photon                 |   |
|   | Absorbing<br>energy   |                                 |  |   |

lifetime of nucleus

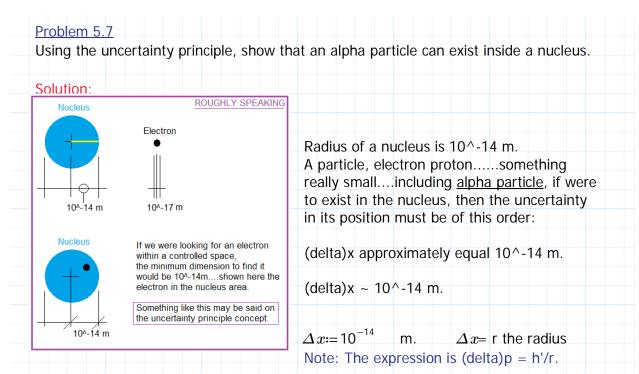
in excited state

e r g

v

Time

| (delta)E                                 | delta)t approxima               | atley equal h'                 |                              |  |
|--|---------------------------------|--------------------------------|------------------------------|--|
| (delta)E t                               | he uncertainty in               | energy approximately           | <u>y equal (</u> h'/delta t) |  |
| $\Delta t = 10^{-1}$                     | <sup>2</sup> seconds            |                                |                              |  |
| $\Delta E \coloneqq \frac{h'}{\Delta t}$ | = 1.055 • 10 <sup>-22</sup>     | J Ans.                         |                              |  |
| Now, lets                                | try to calculate si             | imilalry the uncertain         | ty in frequency.             |  |
| (h)(delta<br>Note its h                  |                                 | <u>v equal (</u> delta)E the u | ncertainty in energy.        |  |
| $\Delta v = \frac{\Delta H}{h}$          | $\dot{L} = 1.592 \cdot 10^{11}$ | Hz. Ans.                       |                              |  |
|  |                                 |                                |                              |  |
|  |                                 |                                |                              |  |
|  |                                 |                                |                              |  |
|  |                                 |                                |                              |  |
|  |                                 |                                |                              |  |
|  |                                 |                                |                              |  |
|  |                                 |                                |                              |  |
|  |                                 |                                |                              |  |
|  |                                 |                                |                              |  |
|  |                                 |                                |                              |  |
|  |                                 |                                |                              |  |
|  |                                 |                                |                              |  |
|  |                                 |                                |                              |  |
|  |                                 |                                |                              |  |



The uncertainty in the momentum of the particle, as shown previouly, must be of the order of:

(delta)p (delta) x = h' (delta)p = h'/(delta)x.

$$\Delta p_{alpha} := \frac{h'}{\Delta x} = 1.055 \cdot 10^{-20}$$
 kg m/s.

Momentum of the particle is at the minimum of this order, i.e. 10<sup>-20</sup>. Notice the order is much lower than (delta)x!

There's no need to open up a Physics textbook on what exacly an alpha particle is, Encyclopedia Britianica is here.

Enclycopedia Britianica: Alpha particle, positively charged particle, identical to the nucleus of the helium-4 atom, spontaneously emitted by some radioactive substances, consisting of two protons and two neutrons bound together, thus having a mass of four units and a positive charge of two. Discovered and named (1899) by Ernest Rutherford, alpha particles were used by him and coworkers in experiments to probe the structure of atoms in thin metallic foils. This work resulted in the first concept of the atom as a tiny planetary system with negatively charged particles (electrons) orbiting around a positively charged nucleus (1909–11). Later, Patrick Blackett bombarded nitrogen with alpha particles, changing it to oxygen, in the first artificially produced nuclear transmutation (1925). Today, alpha particles are produced for use as projectiles in nuclear research by ionization—i.e., by stripping both electrons from helium atoms—and then accelerating the now positively charged particle to high energies.

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Exercise by: K S Bogha. Part IV Wave Packets and The Uncertainty Principle.

Rest mass of an alpha particle is roughly 4 times the mass of a proton.

 $m_{proton} = 1.67 \cdot 10^{-27}$  kg.

 $m_{alpha} := 4 \cdot m_{proton} = 6.68 \cdot 10^{-27}$  kg.

The speed of the alpa particle thus would be? From the momentum relation (delta)p = mv, where v = (delta) p/m.

 $V_{alpha} \coloneqq \frac{\Delta p_{alpha}}{m_{alpha}} = 1.58 \cdot 10^6 \text{ m/s}$ 

The speed does not seem very high...because the order of the speed of light is 10^8, and the alpha particle is at 10^6, just 10^3 difference.

What does this say?

Its in the non-relativistic area. So we use the non-relativistic expression for KE.

 $\mathsf{KE}_{\mathsf{alpha}} \coloneqq \frac{\left(\Delta p_{\mathsf{alpha}}\right)^2}{2 \cdot \mathsf{m}_{\mathsf{alpha}}} = 8.334 \cdot 10^{-15} \text{ J}$ 

KE in eV:

 $KE_{alpha_{eV}} := \frac{KE_{alpha}}{eV} = 5.209 \cdot 10^4$ 

 $KE_{alpha eV} = 52.09 \cdot 10^3$  keV. Ans.

How do we go further? To proof the GROUND STATE energy of the helium atom?

Since its the GROUND STATE, we are saying the energy is at its lowest value. Obviously, then one way to proof is that at higher state the alpha particle energy would be greater then  $52.09 \times 10^3$  keV. Such is the case, when energy carried by an alpha-particle emitted by nuclie it is much higher than this value. Now we can say alpha-particles can exist inside a nucleus.

Comment:

"Some of us may say such is not the best case of defence going to court, but such is the subject matter of quantum mechanics" - Karl Bogha.

A rather interesting example to had given the <u>quantum oppurtunity</u> to make that comment.

|                              | n) is confined to a nucleus of radiuis 5x10 <sup>^</sup> -15 m.<br>ible values of the momentum and the kinetic energy      |
|------------------------------|--|
| Solution:                    |  |
| Nucleus<br>neutron OR proton | We read it similar to the previous problem, not<br>exactly the same, but some toughts do lead us to<br>the same equations. |
| 5x 10 ^-15 m 10 ^-17 m       | Review again the previous problem. To get<br>the minimum possible value what needs to be<br>adjusted in the expression     |
| Nucleus                      | (delta)p = h' / (delta)x?  |
| 5x 10 ~ 15 m                 | delta(x) needs to be maximised or set<br>maximum so (delta)p is minimum because h'<br>is the same in the expression.       |

If (delta)x is made small, then delta(k) must get larger, then their multiplication is greater than or equal to 1.

So obviously we cannot make both widths small or both widths large.

## <u>Theory:</u>

The smaller the spatial extent, (delta)x, of a wave packet, the larger the range of wave numbers, (delta)k, in its Fourier decomposition, and vice versa.

The exact statment and proof of the position momentum uncertainty relation: using root-mean square deviation results with

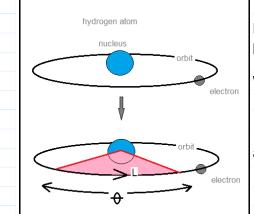
## (delta)x (delta)p > = h'/2.

We know (delta)x (delta)k >=1, h'k = p, thus (delta)x (delta)p >= h'.
Then.....from applying..
...Schrodinger Equations....Schwarz inequality... (delta)x (delta)p > = h'/2.
Maximum uncertainty in the position of the nucleon( neutron or proton) is going to be at either side. So we multiply the radius by 2. Why? see above notes.
(delta)x (delta)p > = h'/2.
2 (delta)x (delta)p > = h'. The momentum is not going to be the same the adjustable is (delta)x, where it is made max resulting in minimum momentum p.

r := 5 · 10<sup>-15</sup> m.  

$$\Delta x$$
:= r  
 $\Delta x_{max} := 2 · \Delta x = 1 · 10^{-14}$  m.  
According to the uncertainty principle, the minimum uncertainty thus in momentum of the particle is  
 $\Delta P_{min} := \frac{h'}{\Delta x_{max}} = 1.055 · 10^{-20}$  kg m/s Ans.  
The momentum cannot be less that this value.  
Continuing with the minium kinetic energy KE:  
 $K E_{min} := \frac{(\Delta P_{min}^2)}{2 \cdot m_{neutron}} = 3.324 \cdot 10^{-14}$  J.  
The answer in eV:  
 $K E_{min_nev} := \frac{K E_{min}}{eV} = 2.077 \cdot 10^{5}$   
 $K E_{min_nev} := 0.2077$  MeV Ans.

| Problem 5.9   |             |                |                 |                |              |
|---------------|-------------|----------------|-----------------|----------------|--------------|
| f the angular | momentum    | of the electro | n in a hydrod   | en atom is kno | wn to be 2h' |
| within 5% ac  | uracy, show | that its andu  | lar position in | a perpendicula | ar plane     |
| annot be spe  |             |                |                 |                |              |



|    | Lets say the figure on the left roughly describes the problem. |
|----|--|
| on | We have linear, straight line, momentum:                       |
|    | (delta)x (delta)p = h'   |
|    | and we got angular momentum:                                   |

(delta) angle theta (delta) angular momentum L = h'.

Uncertainty principle for angular momentum:

(delta)theta (delta)L approximately equal h'

L = 2h' with 5% accuracy

(delta)L = 2h' (5/100) = h'/10

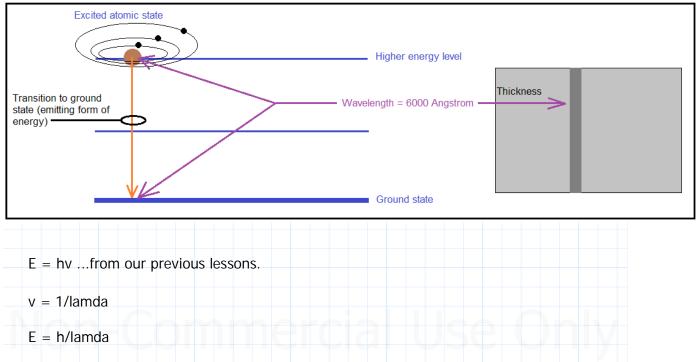
(delta) theta = h' / (delta)L = h' (10/h')

= 10 radians

1 radian = 2 pi, the angle obviously cannot be greater than 1 radian, we have 10 radian!

Conclusion: With L = 2h' the angular position in a perpendicular plane cannot be determined. Ans.

| U U                | etime of an excited atomic state is 10 <sup>^</sup> -8s.                   |
|--------------------|--|
|                    | th of the spectral line associated with the transition from                |
| this state to the  | e ground state is 6000 Angstrom, estimate the width of this line.          |
| Solution:          |  |
| In this evercise   | I consider this an informative problem.                                    |
|                    | he outcome from the subject matter studied through the QM                  |
|                    | ecause the contents is quite deep with many lengthy derivations.           |
|                    | turn around and do this short solution to the question for the line        |
| width, its a littl | e accomplishment!  |
| The width of th    | e spectral line is what we seek. This is saying if there was an experiment |
|                    | spectral line, and that line's width is what we seek.                      |
| Not easy!          |  |
| Lets see what N    | Mahesh the Physicst-Lecturer presented for the solution.                   |
| Pough skotch r     | provided in figure below it tries to capture the problem. May not be       |
|                    | correct sketch on identifying the thickness as shown. You are most         |
|                    | rect it. Check the internetspectroscopy!                                   |

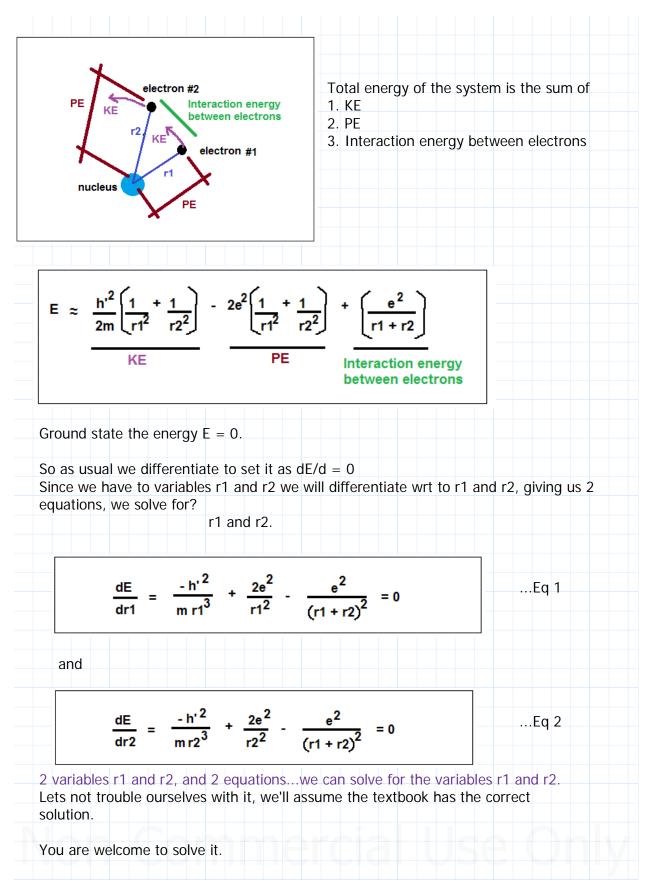


| $(E)_u = Hc/$   | (lamda) d_lamda  |
|---|--|
| becomes   |  |
| E d_Energy  | = hc (1/lamda^2) d_lamda   |
| d_Energy a  | nd D-lamda the differential delta  |
|   |  |
| ∆E =  | $\frac{(hc)}{(A)^2} \triangle A$ This the expression intended to show, was not easy writting it as shown above.              |
| Clearly we  | see the (delta) lamda term represents the width of the spectral line of concer   |
| So that is w  | hat we need to solve for. We have seen similar expression in past  |
| solution of   | differential equation course, here its one of those examples.  |
| The uncerta   | ainty principle of concern here is   |
| (delta)E (de  | elta)t = h'  |
| substitute f  | or (delta)E  |
| (hc/lamda^  | 2) (delta) lamda(delta)t = h'  |
|   | a = $(lamda^2) / (2 pi c (delta)t)$  |
| (delta)lamd   |  |
| $\lambda \coloneqq 6000 \cdot A$  | $ngstrom  \lambda = (6 \cdot 10^{-7}) m$   |
| $\lambda \coloneqq 6000 \cdot \mathbf{A}$ $\lambda \coloneqq 6 \cdot 10^{-7}$                     | m  |
| $\lambda \coloneqq 6000 \cdot \mathbf{A}$ $\lambda \coloneqq 6 \cdot 10^{-7}$                     | m  |
| $\lambda := 6000 \cdot A$ $\lambda := 6 \cdot 10^{-7}$ $\Delta t := 10^{-8}$                      | m  |
| $\lambda := 6000 \cdot A$ $\lambda := 6 \cdot 10^{-7}$ $\Delta t := 10^{-8}$ $c = 3 \cdot 10^{8}$ | m<br>s<br>m/s  |
| $\lambda := 6000 \cdot A$ $\lambda := 6 \cdot 10^{-7}$ $\Delta t := 10^{-8}$ $c = 3 \cdot 10^{8}$ | m<br>s<br>m/s<br>2)<br>$\mathbf{c} \cdot \Delta t$ = 1.91 · 10 <sup>-14</sup> m Ans.<br>This is an extremely thin thickness. |
| $\lambda := 6000 \cdot A$ $\lambda := 6 \cdot 10^{-7}$ $\Delta t := 10^{-8}$ $c = 3 \cdot 10^{8}$ | m<br>s<br>m/s<br>$\frac{2}{1} = 1.91 \cdot 10^{-14}$ m Ans.  |
| $\lambda := 6000 \cdot A$ $\lambda := 6 \cdot 10^{-7}$ $\Delta t := 10^{-8}$ $c = 3 \cdot 10^{8}$ | m<br>s<br>m/s<br>2)<br>$\mathbf{c} \cdot \Delta t$ = 1.91 · 10 <sup>-14</sup> m Ans.<br>This is an extremely thin thickness. |

Problem 5.11 Using the uncertainty principle, estimate the ground state energy of the helium atom. Solution: When we hear Helium we are taught to comapre it to Hydrogen atom, they have some similarities. Helium atom has 2 electrons. electron #2 Uncertainty principle using momentum and radius: r2 p2 ≈ h'/r2 electron #1 (delta) r (delta) p = h'r1 p1≈ h'/r1 (delta) r1 (delta) p1 = h'nucleus (delta) r2 (delta) p2 = h' Kinetic energy of the system, composed of the nucleus and 2 electrons: KE equal approximately  $(h'^{2}/2m) ((1/r1^{2}) + (1/r2^{2}))$ m the electron mass in the expression above. PE of the interaction of 'electrons' with the 'nucleus of charge 2e' since we have 2 electrons is: PE equal approximately  $-2 e^2$  ((1/r1) + (1/r2)) .....from Physics 1st year college/university. Now, the ORDER of separation 'between the electrons' is the 'sum of the order of r for each electron': (r1 + r2).....uncertainty principle. The interaction energy between the electrons is approximately:  $(e^2) / (r1 + r2).$ 

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$$r_{1} = r_{2} = \frac{4}{7} \left( \frac{h^{2}}{me^{2}} \right)$$
Substituting back r1 and r2 into the euqation for E, we have
$$E = -\frac{49}{16} \left( \frac{me^{4}}{h^{2}} \right)$$
Solve for E above by substitution of values.
$$m_{electron} = 9.1 \cdot 10^{-31} \qquad h' = 1.05 \cdot 10^{-34} \qquad eV = 1.6 \cdot 10^{-19}$$

$$E := \left( \frac{-49}{16} \right) \cdot \left( \frac{(m_{electron}) \cdot (e^{4})}{(h^{2})} \right) = -1.657 \cdot 10^{-37} \text{ J}$$
In eV:
$$E_{eV} := \frac{E}{eV} = -1.035 \cdot 10^{-18}$$

$$E_{eV} := -10.35 \quad eV \text{ Ans.}$$

End of Chapter 5 exercises from QM for UG by Mahesh C. Jain.

The next chapter is on the theory side of Schroedinger Equation. No solved problems here.

## **Recommended textbook on Schroedinger Equation:**

Quantum Mechanics DeMystified: A Self-Teaching Guide. David McMahon. 2006. McGrawHill. Chapters 1 and 2.